

# Indeterminacy of Quantum Geometry

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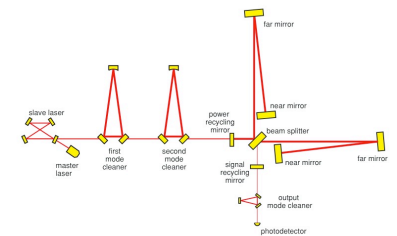


## Wave theory of quantum geometry

- Standard quantum field theory: geometry is classical, fields are quantized, modes are independent, "all physics is local"
- Gravitation theory suggests holographic quantum geometry: UV cutoff at the Planck length, 3D states and correlations are spatially nonlocal
- **New holographic phenomenology: relative spacetime positions of events are indeterminate, defined to precision of information propagation and measurement allowed by capacity of Planck wavelength radiation**
- spacetime events and paths complementary to Planck wavelength radiation (classical event~ particle, path~ray)
- **Theory of spacetime indeterminacy based on wave optics: wavefunctions represent event positions**

# Direct measurement of **quantum geometry fluctuations**

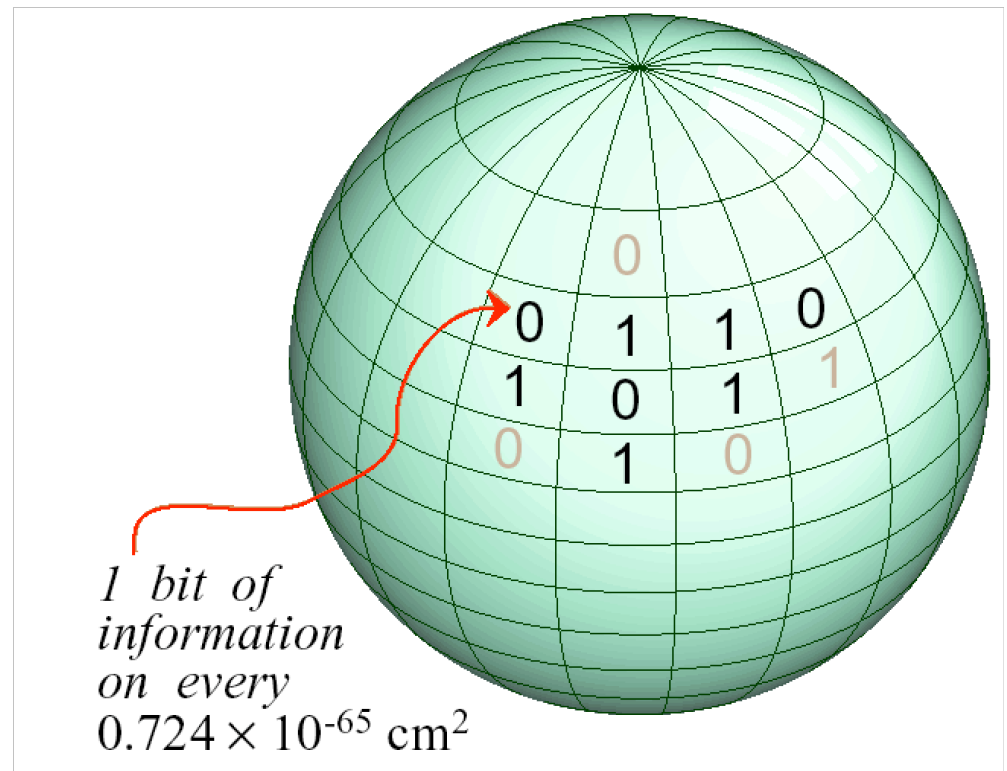
- Experimental approach to small lengths/large energies has been the high energy frontier-- accelerators
- Better macroscopic position measurements by interferometers designed for gravitational wave detection (narrow wavefunction of macroscopic mass position: small Heisenberg uncertainty)
- Quantum wave geometry predicts a new detectable effect: **"holographic noise"**
- black hole evaporation physics--- in the lab
- Spectrum and distinctive spatial character of the noise is predicted with no parameters
- An experimental program is motivated
- CJH: [arXiv:0806.0665](https://arxiv.org/abs/0806.0665), also PRD D 77, 104031 (2008)



“This is what we found out about Nature’s book keeping system: the data can be written onto a surface, and the pen with which the data are written has a finite size.”

-Gerard 't Hooft

*Everything about the 3D world can be encoded on a 2D surface at Planck resolution (?)*



# Holographic Quantum Geometry: theory

- Black holes: entropy=area/4  $S = A/l_P^2 4 \ln 2$
- Black hole evaporation
- Einstein's equations from heat flow
- Classical GR from surface theory
- Universal covariant entropy bound
- Exact state counts of extremal holes in large D
- AdS/CFT type dualities: N-1 dimensional duals
- All suggest that quantum geometry lives on 2+1 dimensional null surfaces

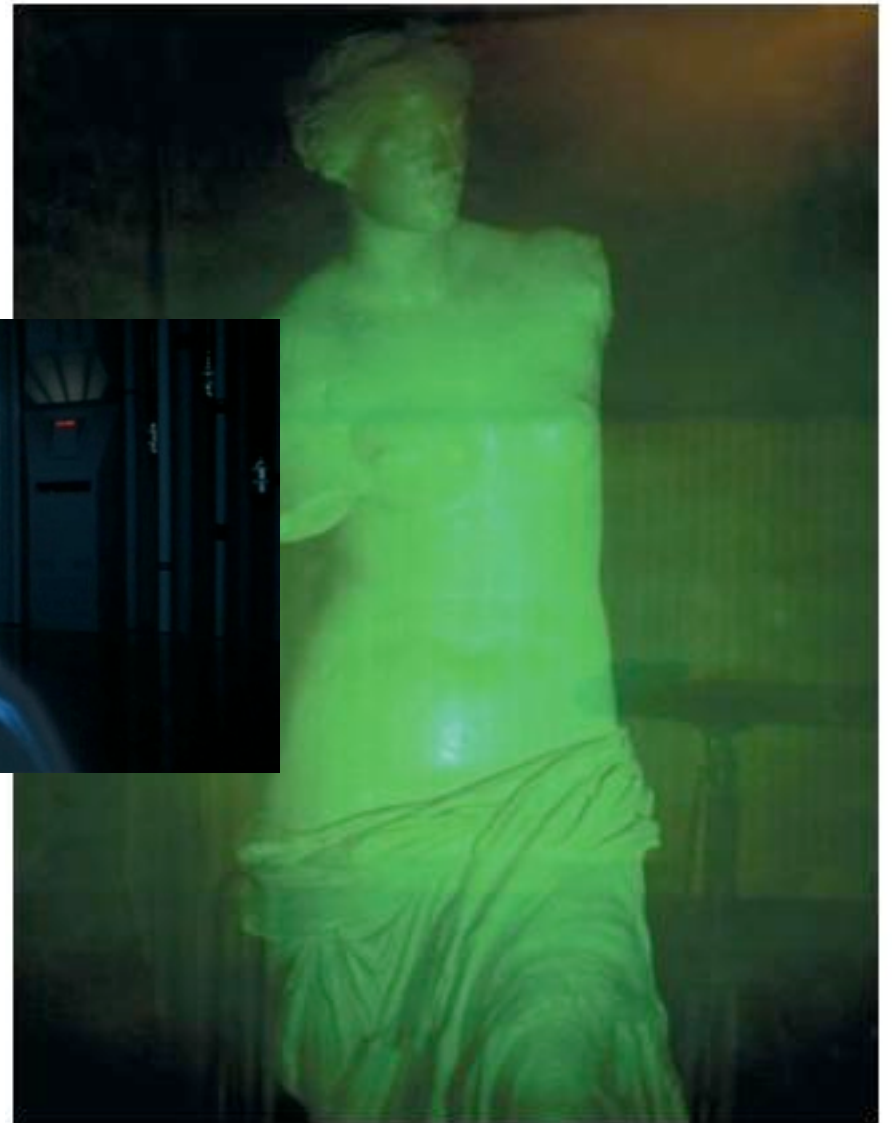
Beckenstein, Hawking, Bardeen et al., 'tHooft, Susskind, Bousso, Srednicki, Jacobson, Padmanabhan

# A holographic world is blurry

*limited information content*



What does it look like  
"from inside"?  
("Flatland" realized with  
waves)





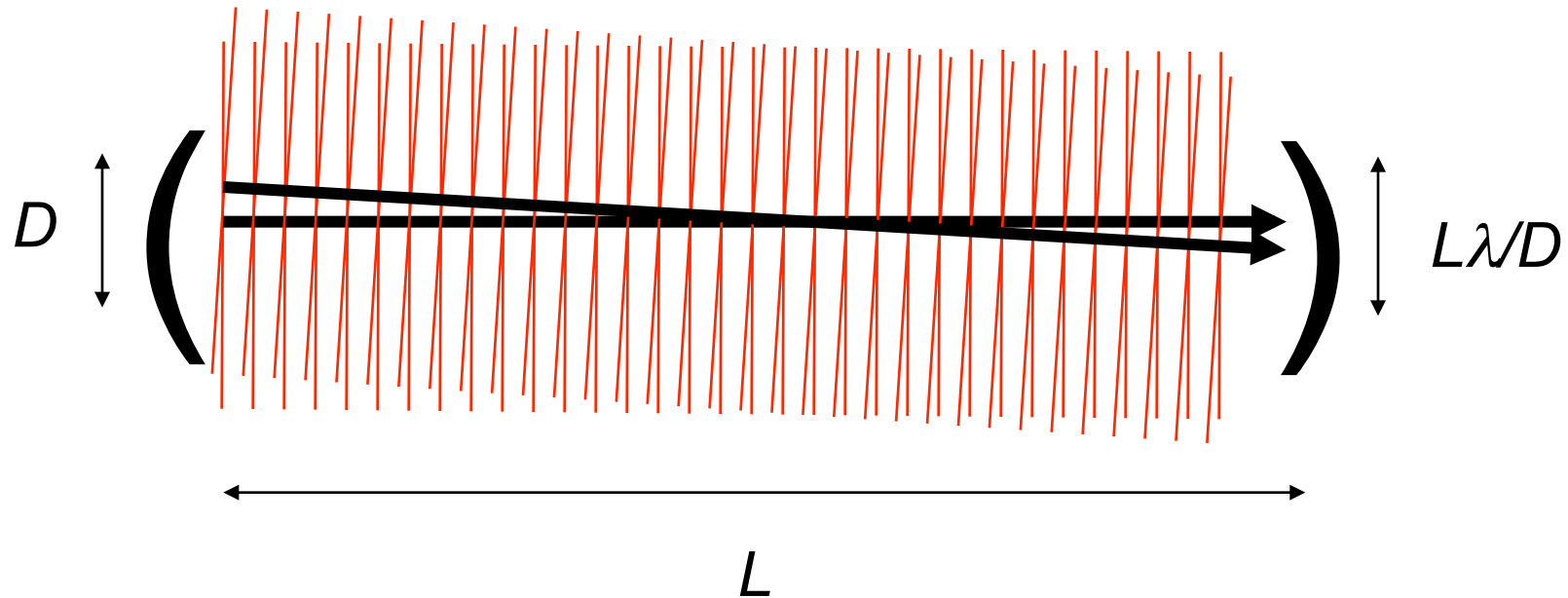
# Holographic Quantum Geometry

- Spacetime is a quantum system
- **Conjecture: the world is formed by Planck wavelength null waves**
- "from inside": transverse quantum fluctuations in position much larger than Planck length

$$l_P = \sqrt{\hbar G_N / c^3} = 1.616 \times 10^{-33} \text{cm}$$



## Ray limit of wave optics: Rayleigh uncertainty



- Aperture  $D$ , wavelength  $\lambda$  : angular resolution  $\lambda/D$
- Size of diffraction spot at distance  $L$ :  $L\lambda/D$
- Endpoints of a ray can be anywhere in aperture, spot
- path is determined imprecisely by waves
- Minimum uncertainty at given  $L$  when aperture size = spot size, or

$$D = \sqrt{\lambda L}$$

## The case of a real hologram

- For optical light and a distance of about a meter,

$$D = \sqrt{\lambda L}$$

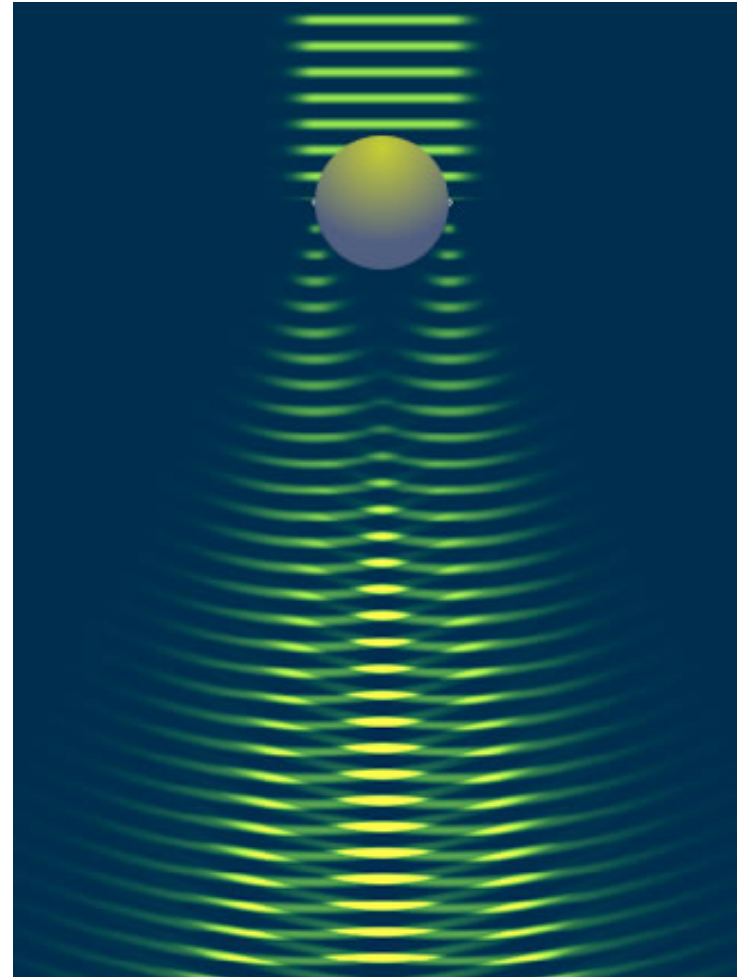
is about a millimeter

- Larger aperture gives sharper image but then photon paths and arrival positions cannot be measured so well
- If you "lived inside" a hologram, you could tell by measuring the blurring/indeterminacy



# Wave Theory of Spacetime Indeterminacy

- Adapt van Cittert-Zernike theory of transverse correlation in wave optics
- theory of “fuzzy event positions”
- Complex amplitude=wavefunction
- Complex correlation=quantum correlation
- Intensity=probability
- Set wavelength to match holographic degrees of freedom



## Uncertainty of transverse position

***Spacetime events are not points but wavefunctions.***

*Transverse to a null trajectory at separation  $L$  events are Fourier transforms of each other and have standard deviations of transverse position related by:*

$$\sigma' \sigma = \lambda L$$

***For macroscopic  $L$  the “uncertainty” is much larger than the wavelength***

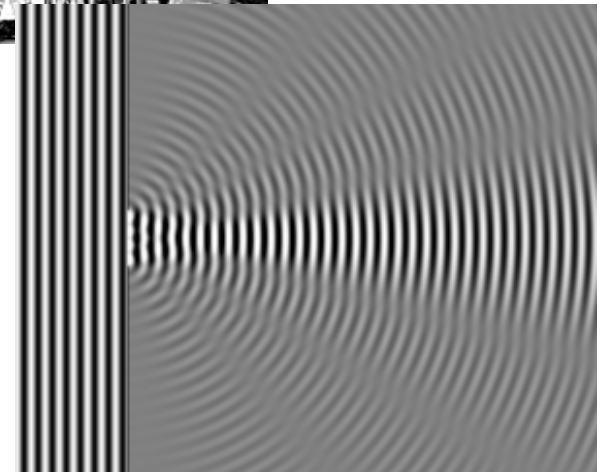
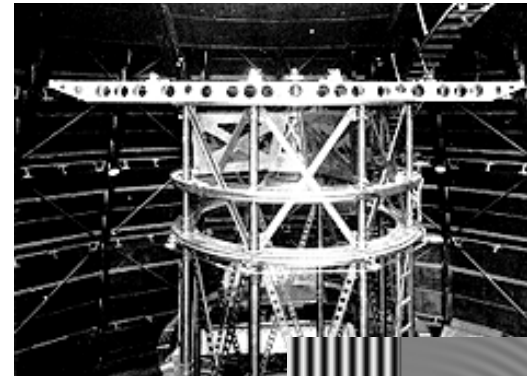
Controlled covariant theory based on wave optics: CJH, arXiv: **0806.0665**

## Familiar examples from the world of optics

- Hanbury Brown-Twiss interferometry: correlation of intensity from distant star in widely separated apertures
- Michelson stellar interferometer: fringes from star
- Diffraction in the lab: shadow of plane wave cast by edge or aperture

All display similar optical examples of wave phenomena much larger than the waves,

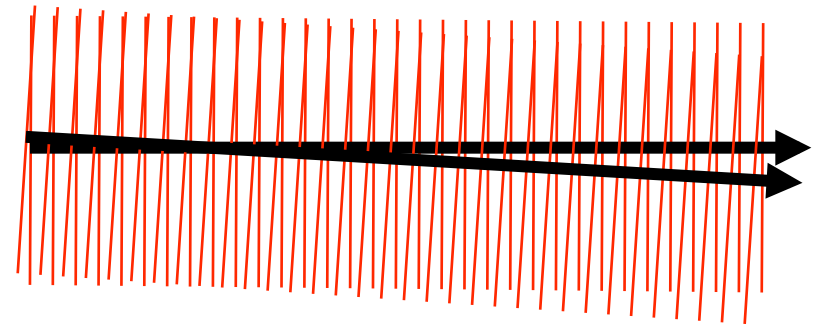
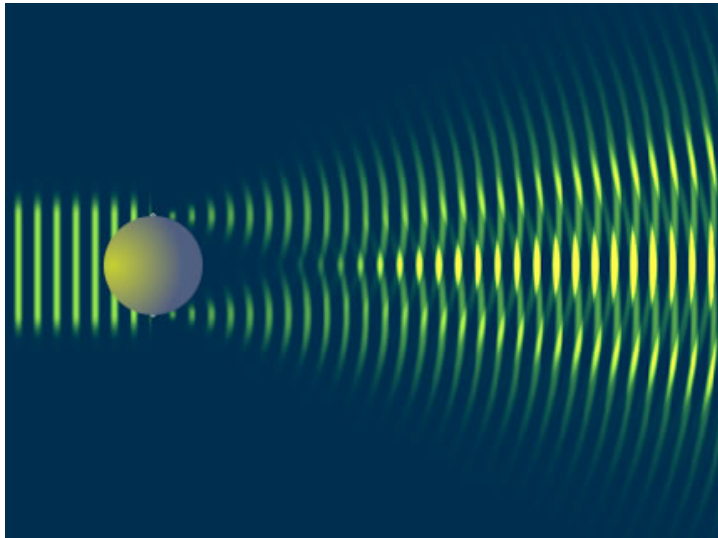
$$\sigma' \sigma = \lambda L$$



Set fundamental wavelength= Planck length

- Agrees with degrees of freedom, minimum wavelength from gravitation theory
- black hole evaporation, unitary behavior of gravitational systems
- Sets absolute normalization with no parameters
- No gravity: flat wavefronts, flat space

## Indeterminacy of a Planckian path: diffraction limit of "Planck wavelength telescope"



- spacetime metric defined by paths between events
- Complementarity: path  $\sim$  ray approximation to waves
- Events on worldline  $\sim$  particle interactions with Planck wavelength radiation
- Transverse wavefunction of events displays indeterminacy formally identical to optical wave correlations
- Indeterminacy of geometry reflects limited information content of wave model



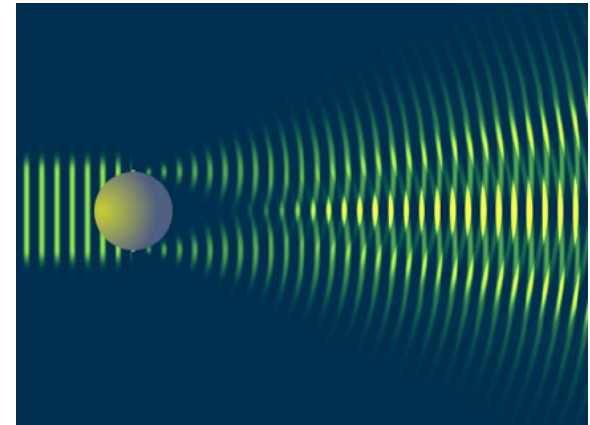
## *holographic approach to the classical limit*

- **Angles** are indeterminate at the Planck scale, and become better defined at larger separations:

$$\Delta\theta(L) = (l_P/L)^{1/2}$$

- But uncertainty in **relative transverse position increases** at larger separations:

$$\Delta x_{\perp}^2 > l_P L$$



- Not the classical limit of field theory
- Indeterminacy and nonlocality persist to macroscopic scales

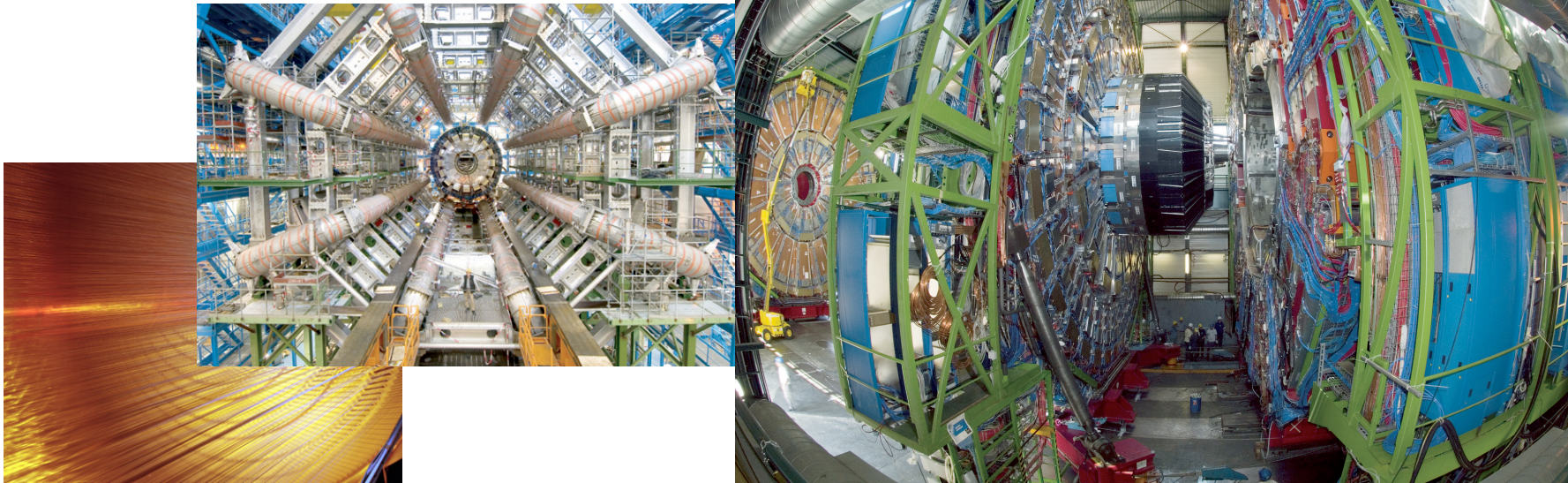
# Fluctuations in quantum geometry

- Distant spacetime is only defined insofar as it can be measured locally using Planck radiation
- Distant events are fuzzy objects, not points
- Endpoints of trajectories (interaction events) are uncertain
- Indeterminacy of worldlines leads to fluctuations in measured quantities
- Statistical predictions not sensitive to model details: direct measure of fundamental information bounds

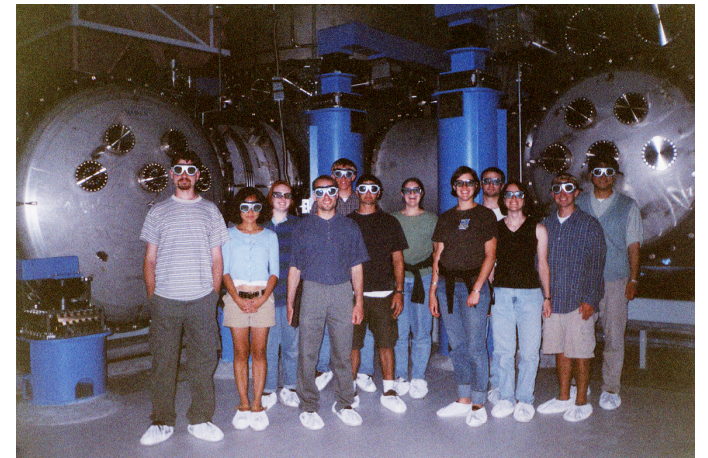


What is the best microscope for measuring quantum geometry?

CERN/FNAL:  $\text{TeV}^{-1} \sim 10^{-18}$  m



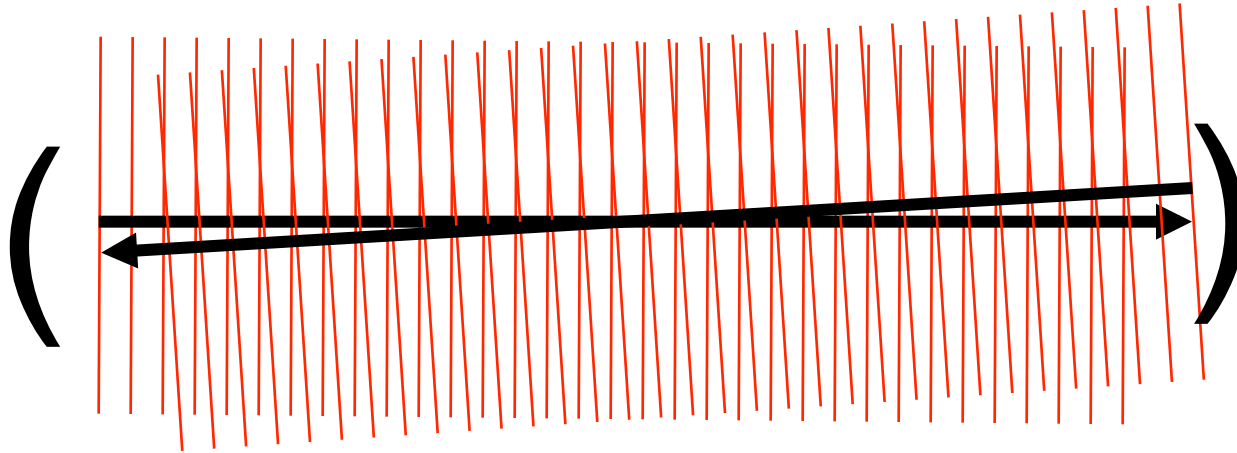
LIGO/GEO:  $\sim 10^{-19}$  m  
over  $\sim 10^3$  m baseline



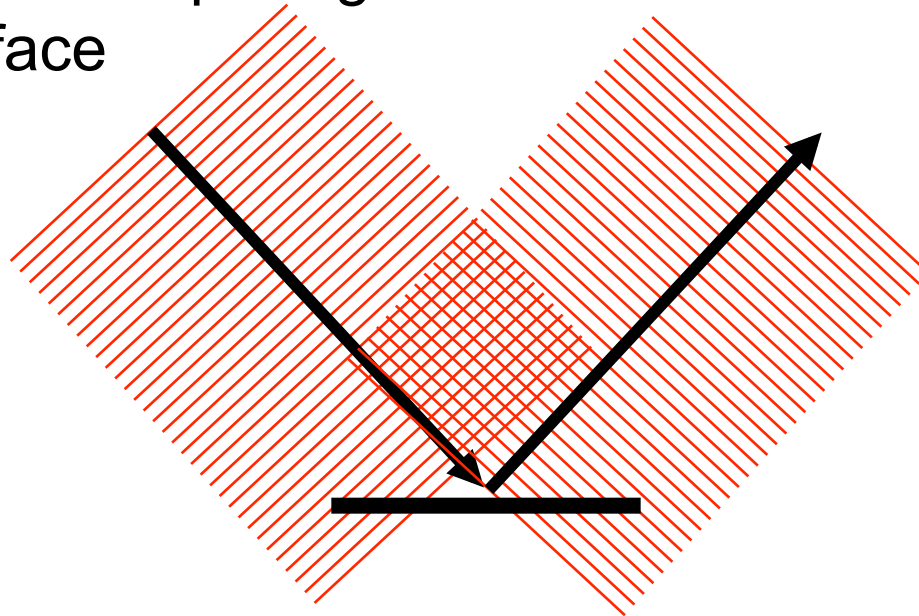
# Interferometers as Planck telescopes

- Nonlocality: precise relative positions at km scales
- Fractional precision: angle  $< 10^{-20}$ , > "halfway to Planck"
- Transverse position measured in some configurations
- Precision: like two collisions at LHC localized at exactly the same place after a complete circuit
- Proof masses have narrow position wavefunction, measure spacetime wavefunction
- Detect holographic blurring: noise in signal stream

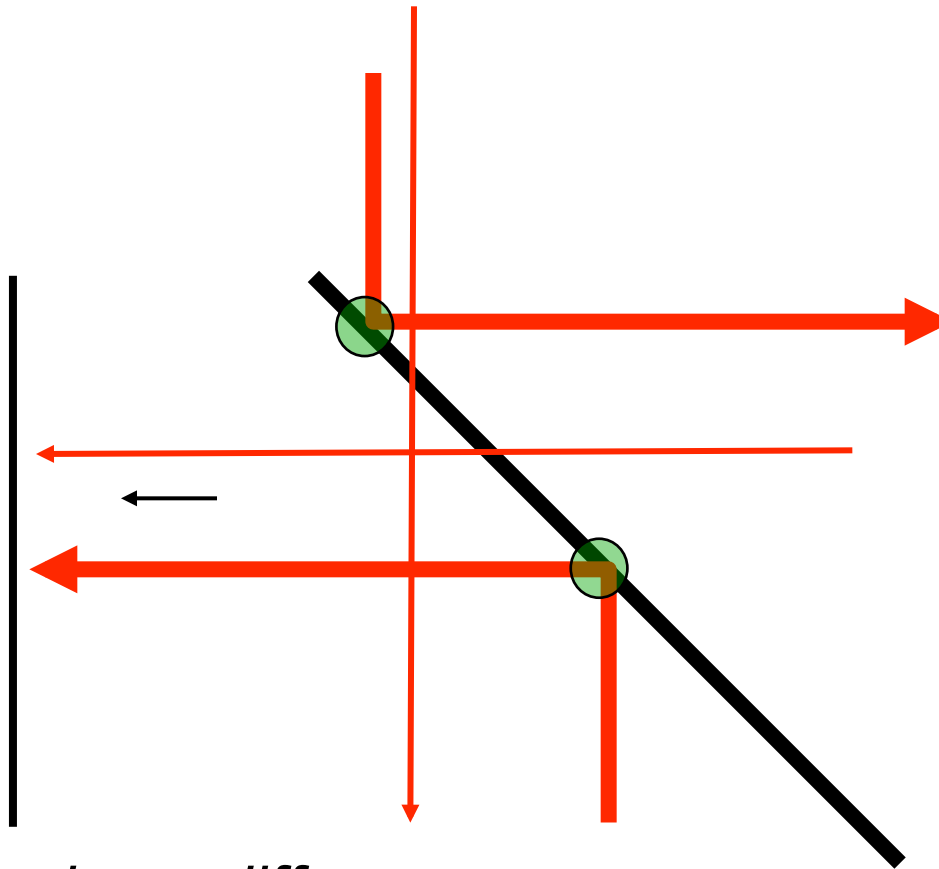
Normal incidence optics: phase signal does not record the transverse position of a surface



- But phase of beam-split signal is sensitive to transverse position of surface



# Beamsplitter and signal in Michelson interferometer



*Signal: random phase difference of reflection events from indeterminate position difference of beamsplitter at the two events*

*Beamsplitter: reflection events at two times separated by  $L=2L_0$*

# Measurement of transverse position of beamsplitter

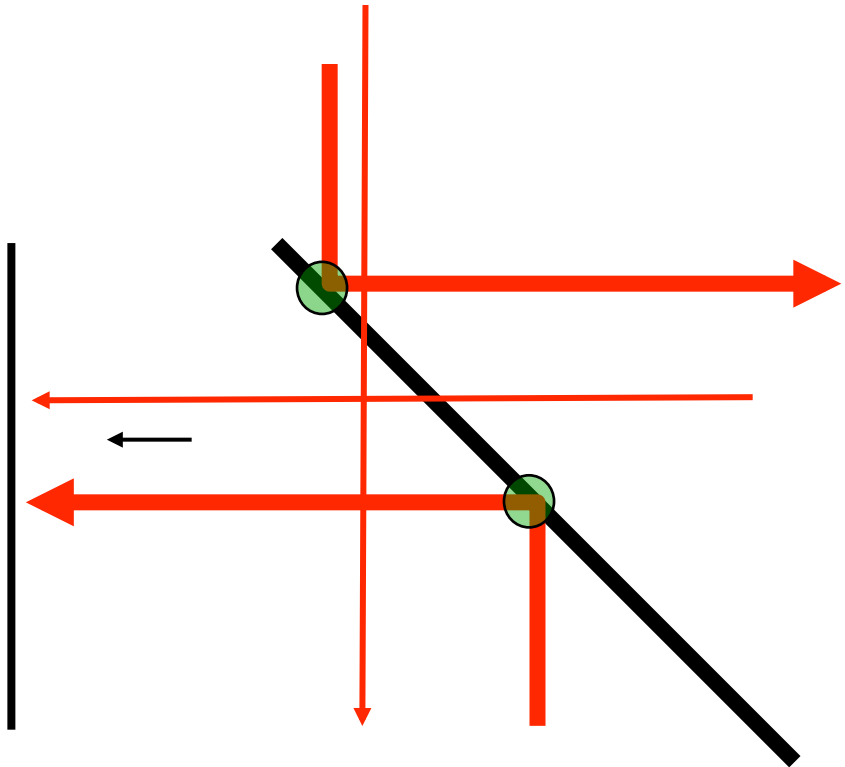
- Positions of reflection events have transverse uncertainties

$$\sigma' \sigma = \lambda L$$

- Independent samplings accumulate signal phase uncertainty
- apparent arm length difference is a random variable, with variance:

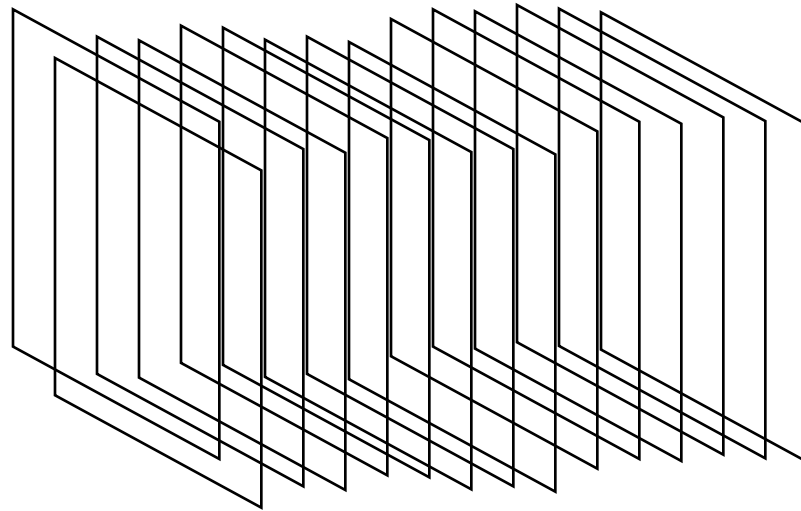
$$\Delta L_0^2 = \sigma^2 + \sigma'^2 = 2\sigma^2 = 4l_P L_0$$

**this is a new effect predicted with no parameters**



## “continuous measurement” ~ transverse Planck random walk

- A highly interactive system mimics a random walk transverse to null direction of propagation (normal measurement): a Planck length per Planck time
- Walk in the position of bodies (e.g., beamsplitter) and not propagating radiation itself





# Power Spectral Density of Fluctuations

Uncertainty in angle  $\sim$  dimensionless metric perturbation

$$\Delta\theta(L) = (l_P/L)^{1/2}$$

*$\sim$  metric shear fluctuations with flat power spectral density*

$$h_H^2 \simeq L\Delta\theta^2 \approx t_P$$

$h_H^2$  = mean square perturbation per frequency interval

(prediction with no parameters, Planck length is the only scale)

# Holographic Noise

*Universal **holographic noise** ~ flat power spectral density of metric **shear** perturbations:*

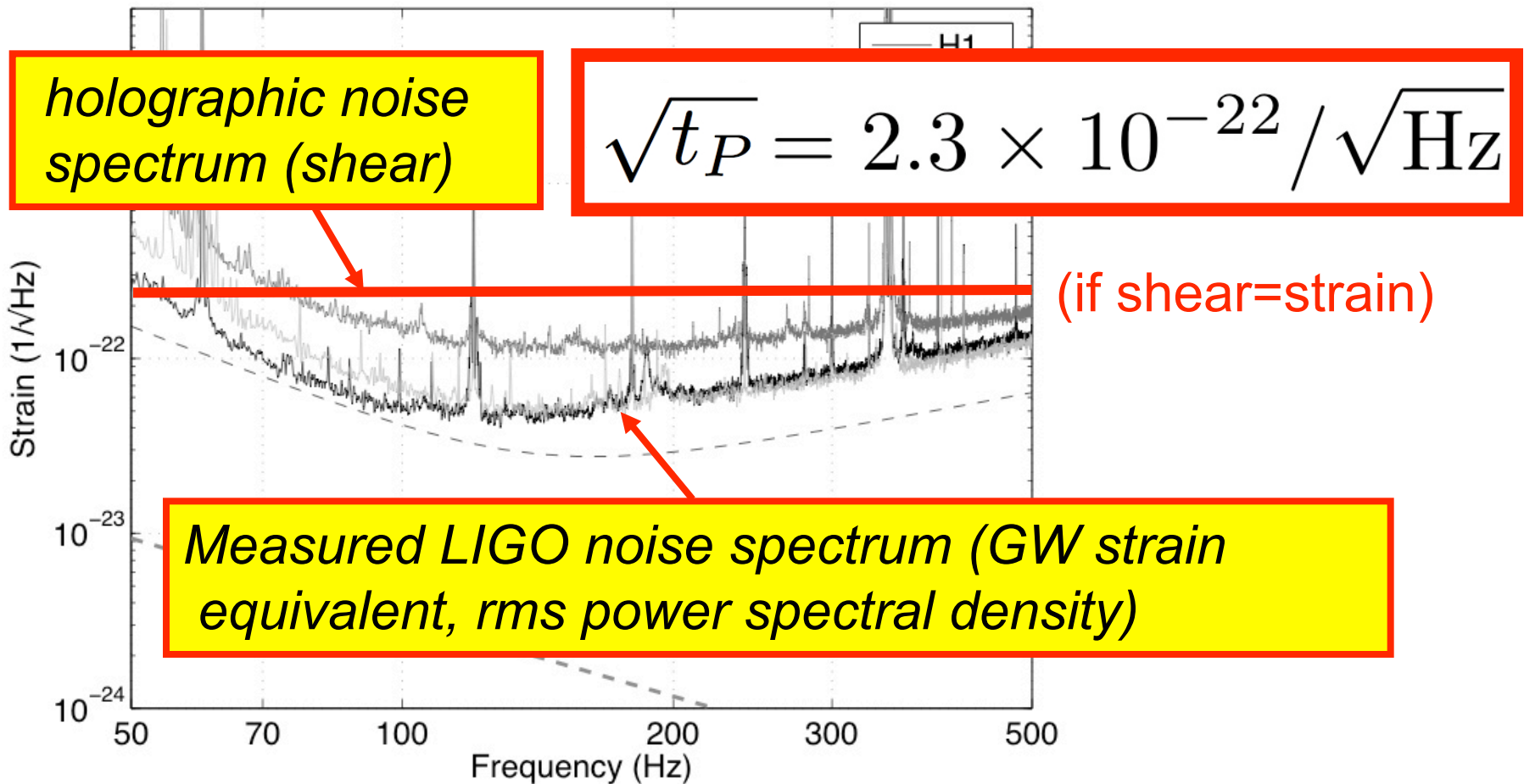
$$h \approx \sqrt{t_P} = 2.3 \times 10^{-22} \text{Hz}^{-1/2}$$

- A property of holographic quantum geometry
- Prediction of spectrum with no parameters
- Prediction of spatial shear character: only detectable in transverse position observables
- Definitively falsifiable
- Contrast with more general range of possible phenomenology (e.g. Amelino-Camelia, et al.)

## Holographic fluctuations do not carry energy or information

- ~ classical gauge mode (flat space, no classical spacetime degrees of freedom excited)
- ~ sampling noise, not thermal noise
- Necessary so the number of distinguishable positions does not exceed holographic bound on Hilbert space dimension
- No curvature
- no strain, just shear
- no detectable effect in a purely radial measurement

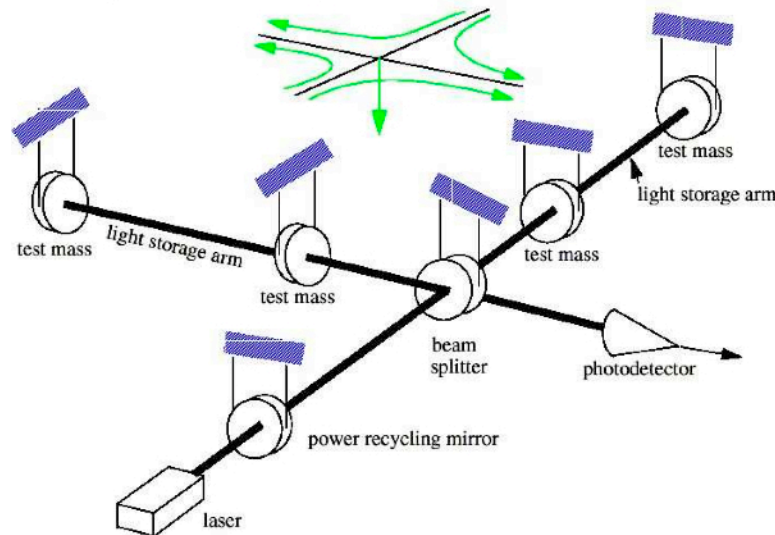
# LIGO noise (astro-ph/0608606)



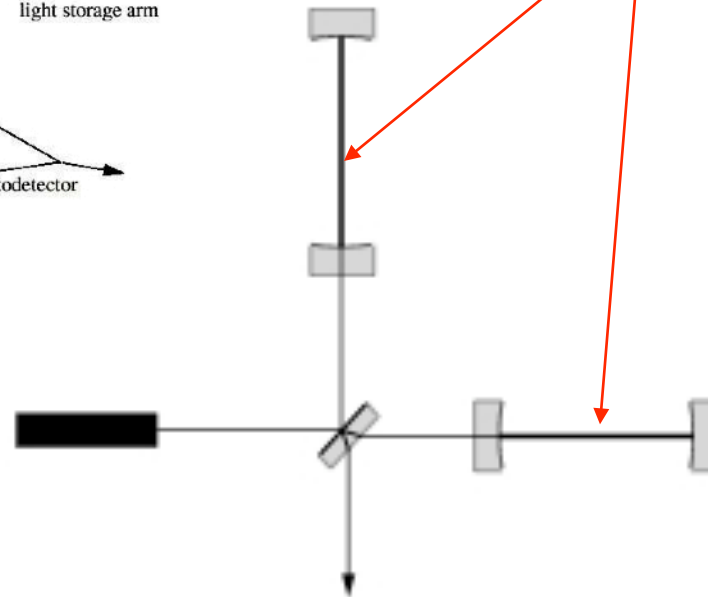
# Why doesn't LIGO detect holographic noise?

- EITHER holographic noise does not exist, OR:
- LIGO layout is not sensitive to transverse displacement noise (relationship of holographic to gravitational wave depends on details of the system layout)

Fig. 1. Schematic layout of a LIGO interferometer.

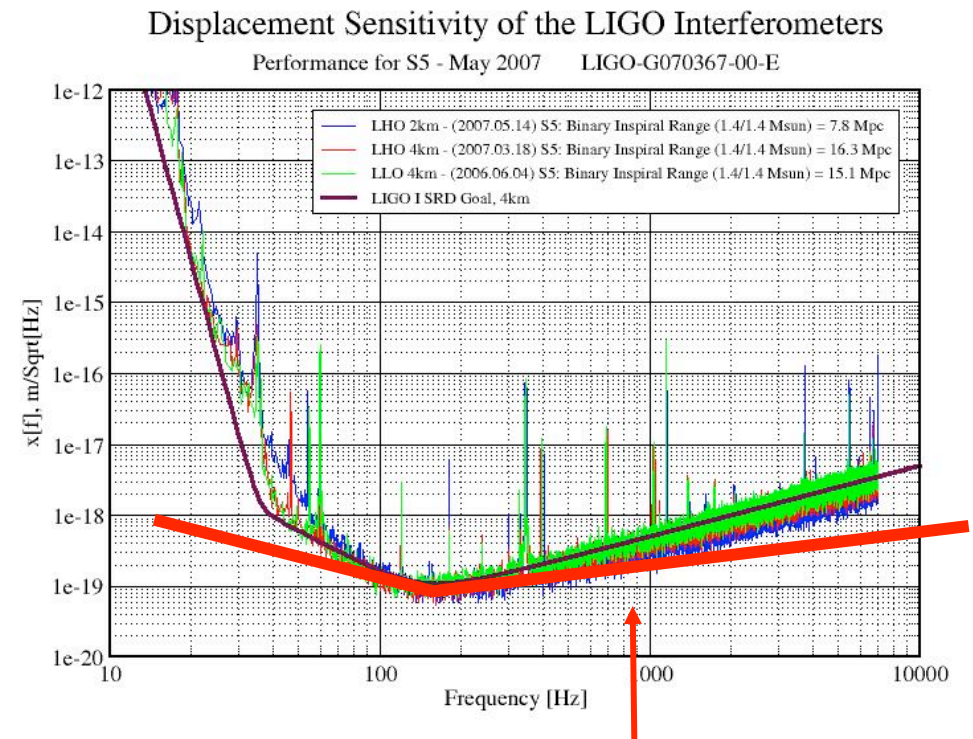


Transverse position measurement is not made in FP cavities



# LIGO S5 run: noise in displacement units

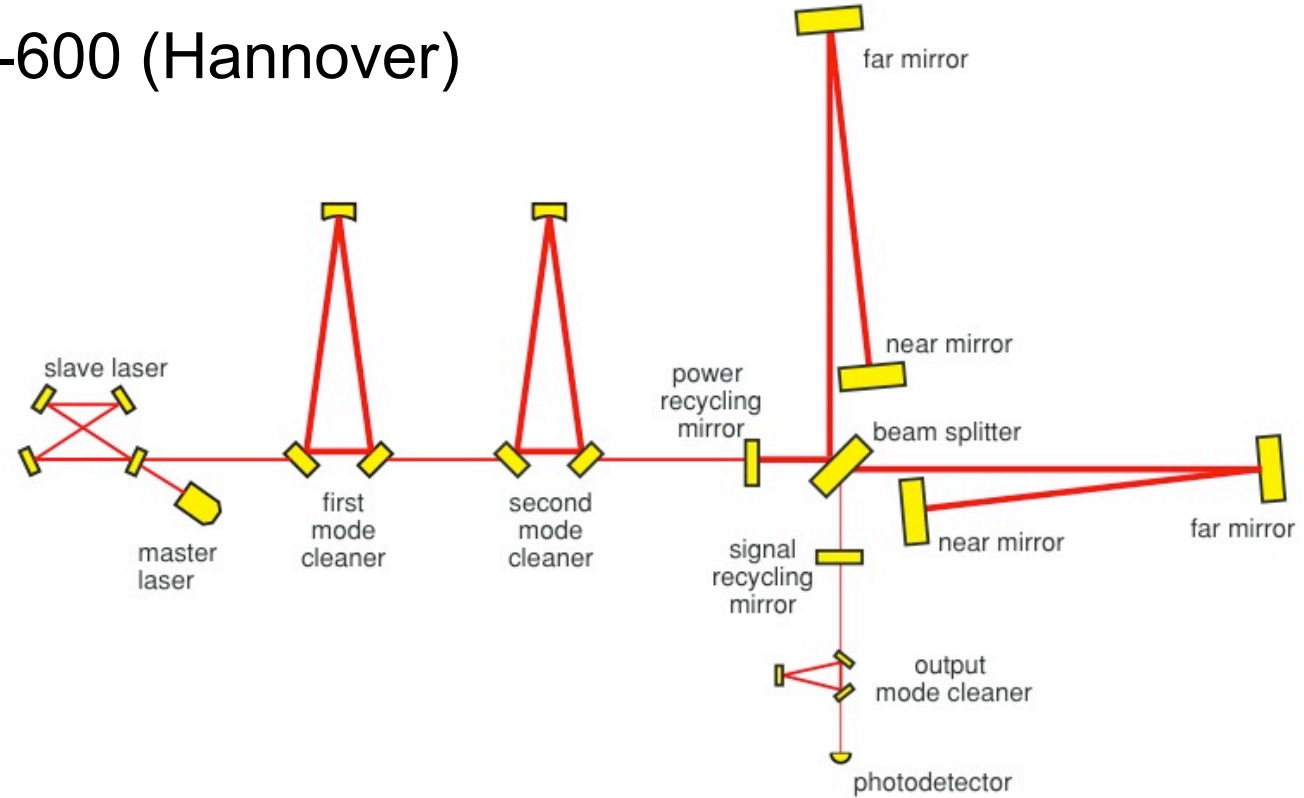
- Allow for lack of holographic noise from FP arm cavities
- In displacement units, estimated holographic noise is below sensitivity of last science run
- Will be detectable with enhanced/advanced LIGO



Rough but zero-parameter estimate of holographic noise in LIGO (displacement units)

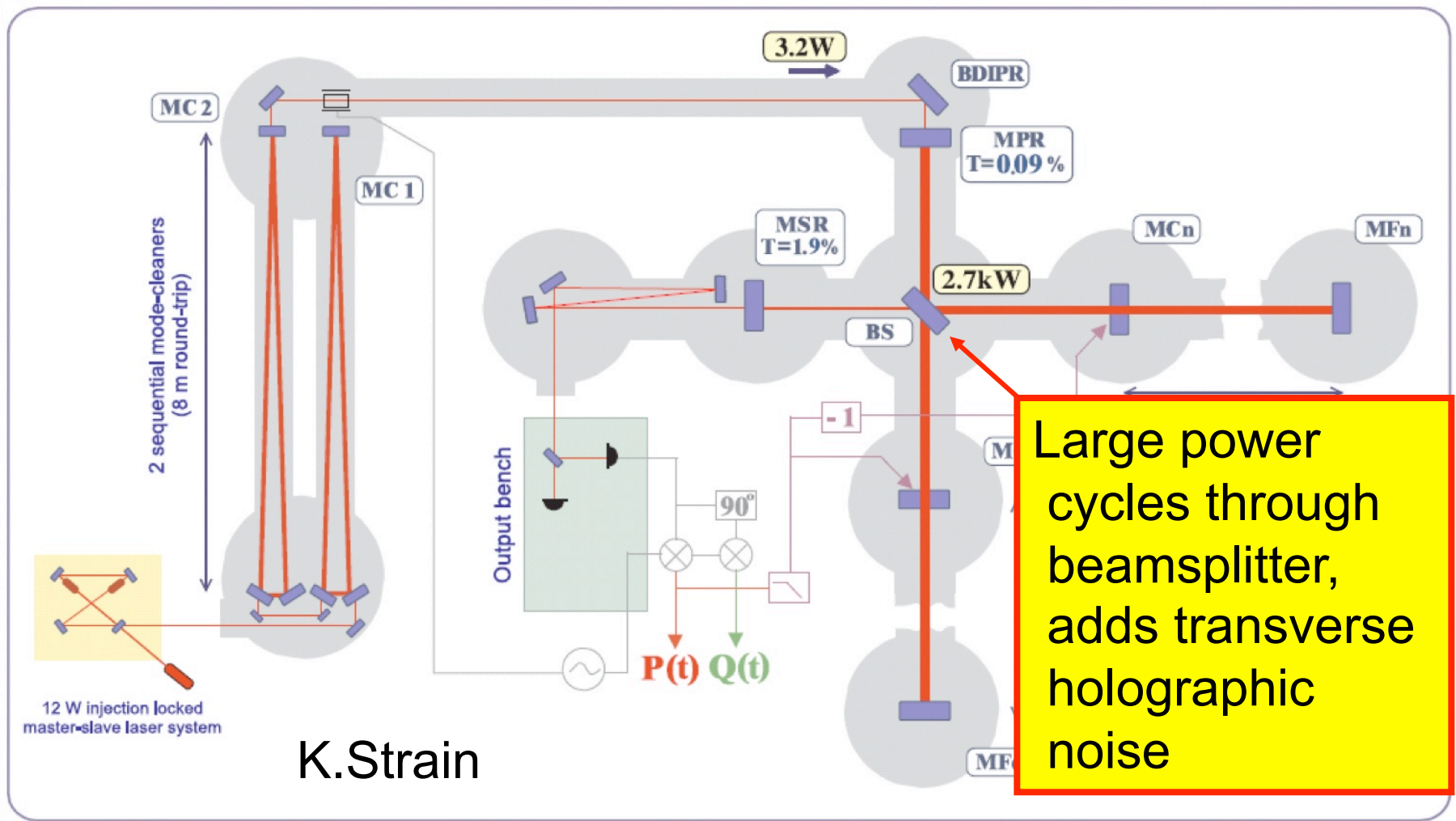
CJH: [arXiv:0806.0665](https://arxiv.org/abs/0806.0665)

# GEO-600 (Hannover)



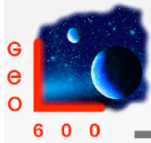


# The GEO600 Interferometer



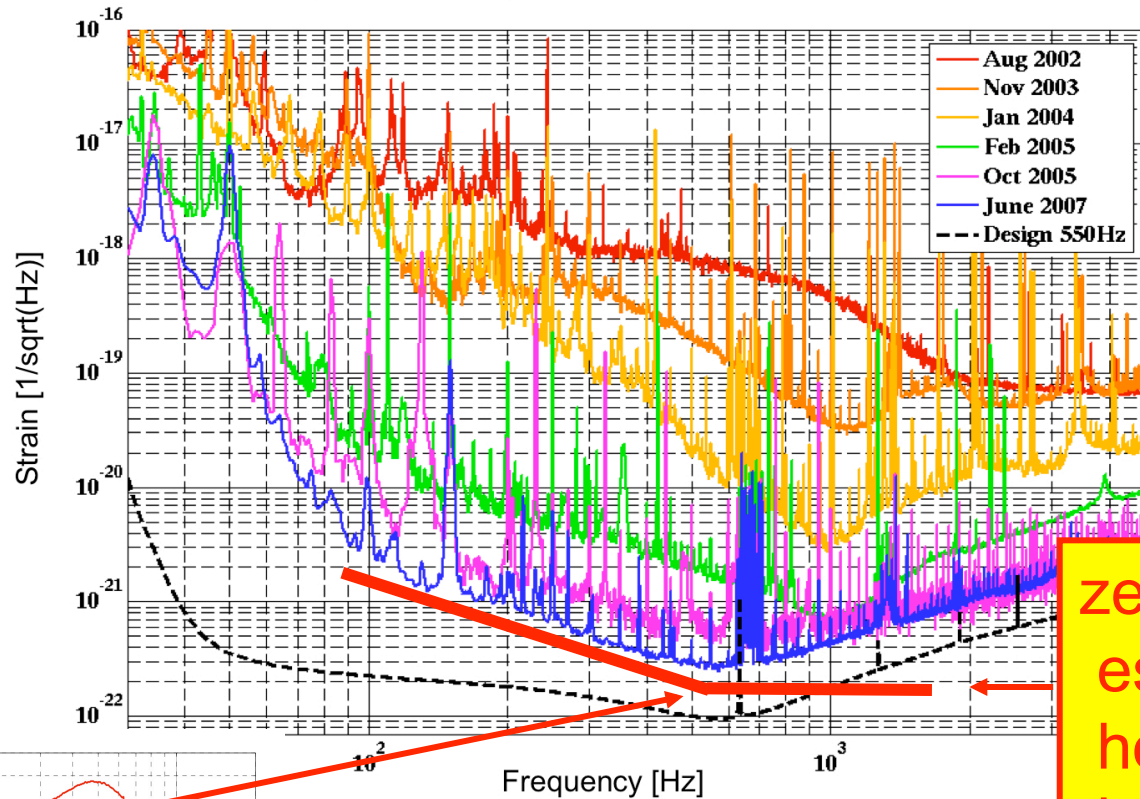


# Noise in GEO600

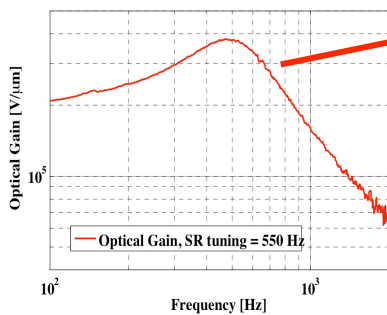


## GEO Sensitivities

K. Strain



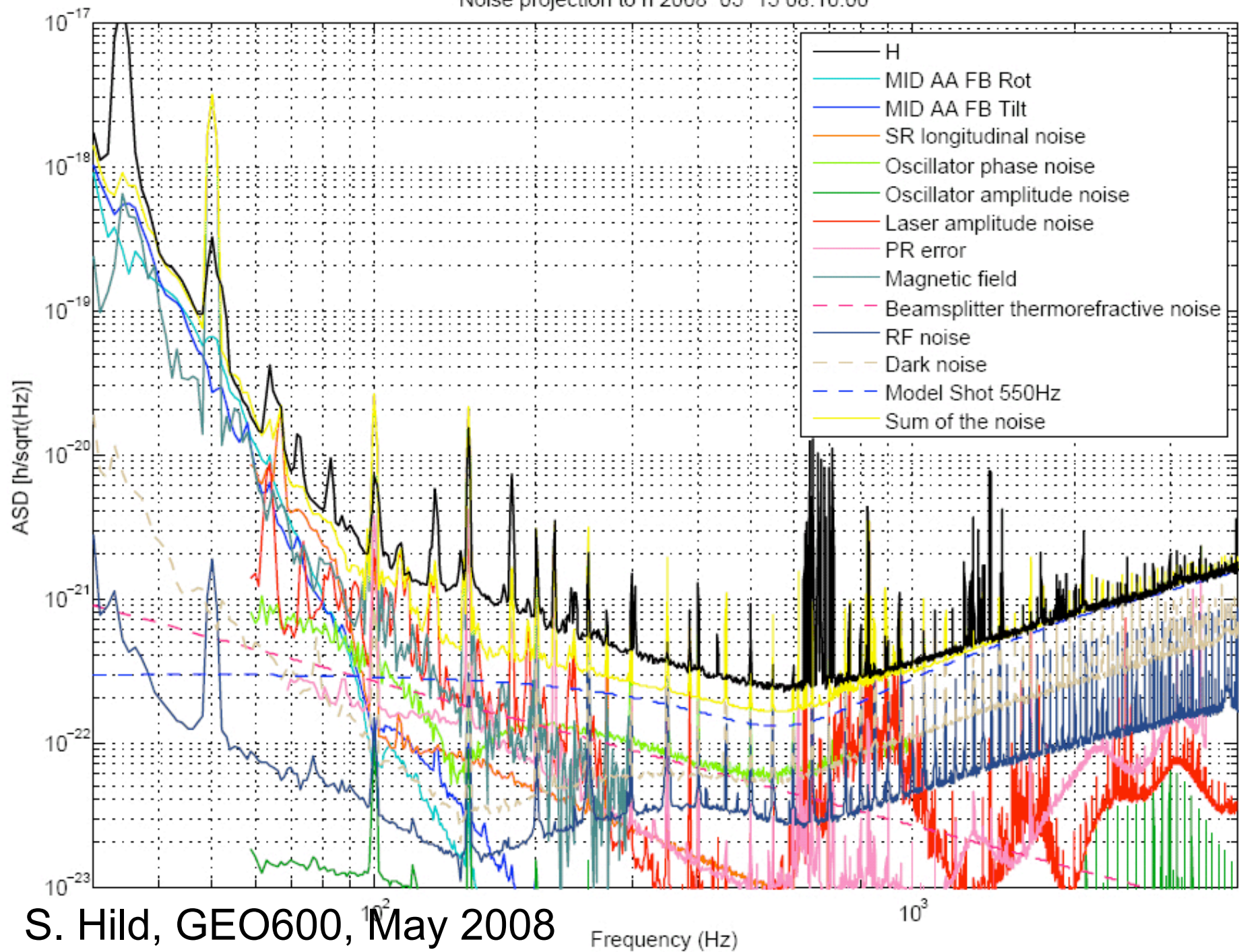
zero-parameter estimate of holographic noise in GEO600 (equivalent strain)



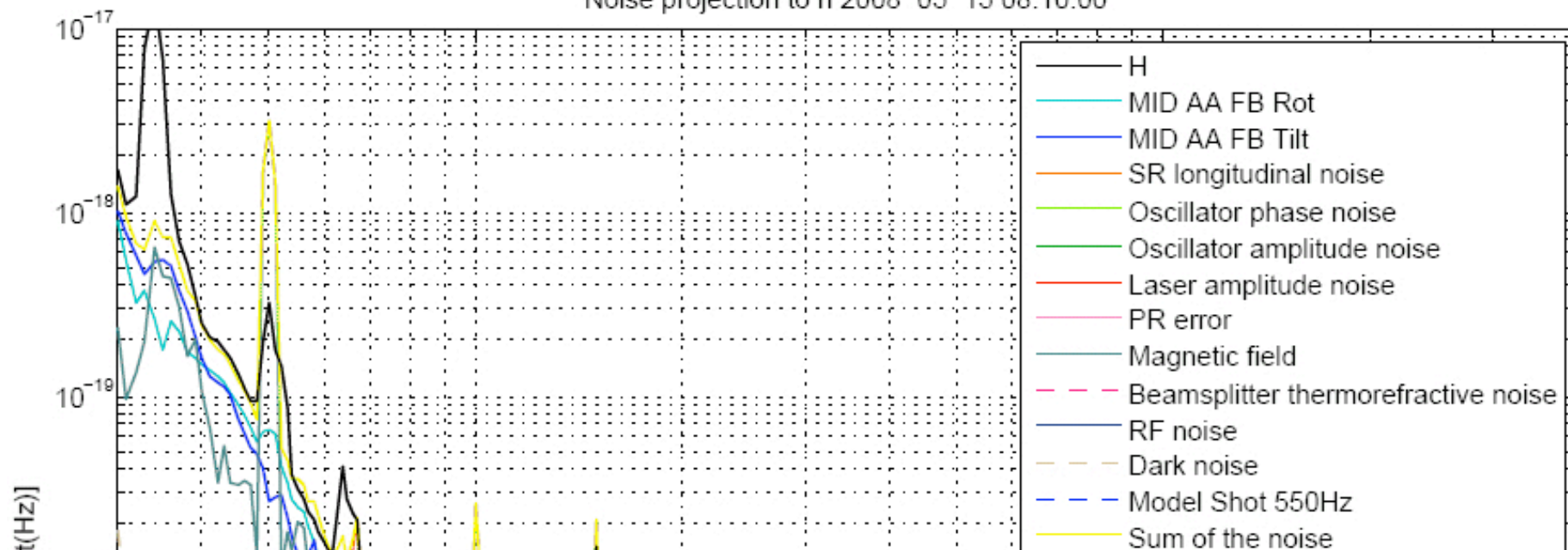
H. Lück, S. Hild, K. Danzmann, K. Strain

CJH: [arXiv:0806.0665](https://arxiv.org/abs/0806.0665)

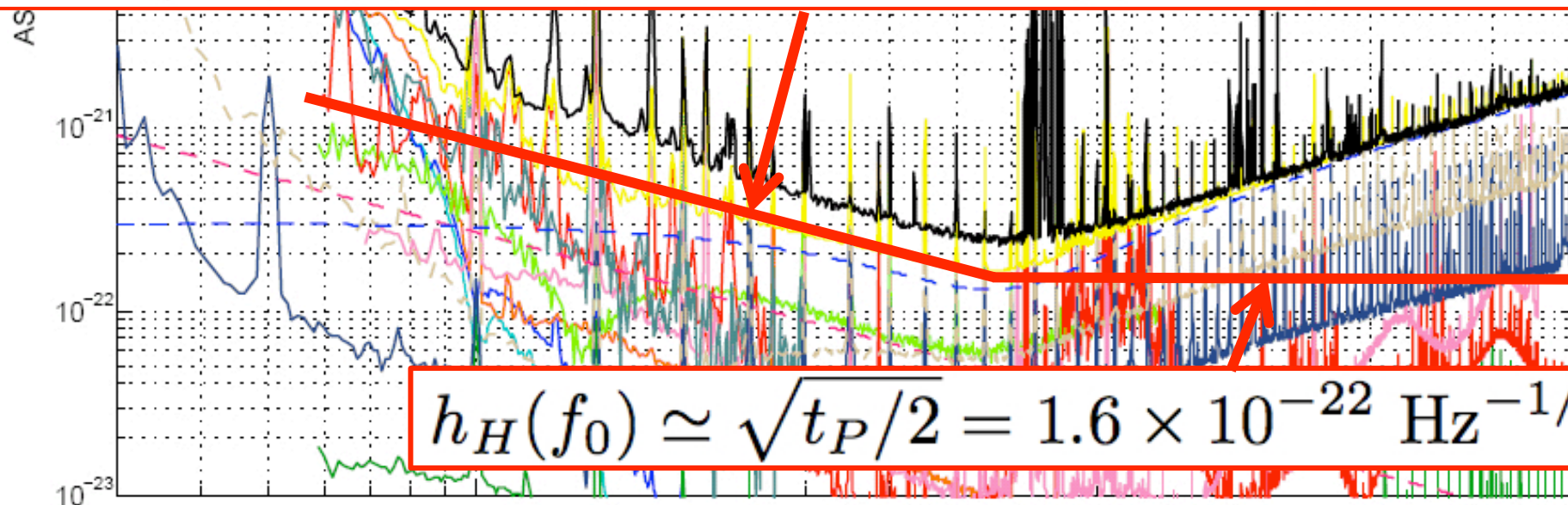
Noise projection to h 2008-05-15 08:10:00



Noise projection to h 2008-05-15 08:10:00



$$h_H \simeq \sqrt{t_P/2} (f/550\text{Hz})^{-1} = 1.6 \times 10^{-22} (f/550\text{Hz})^{-1} \text{ Hz}^{-1/2}$$

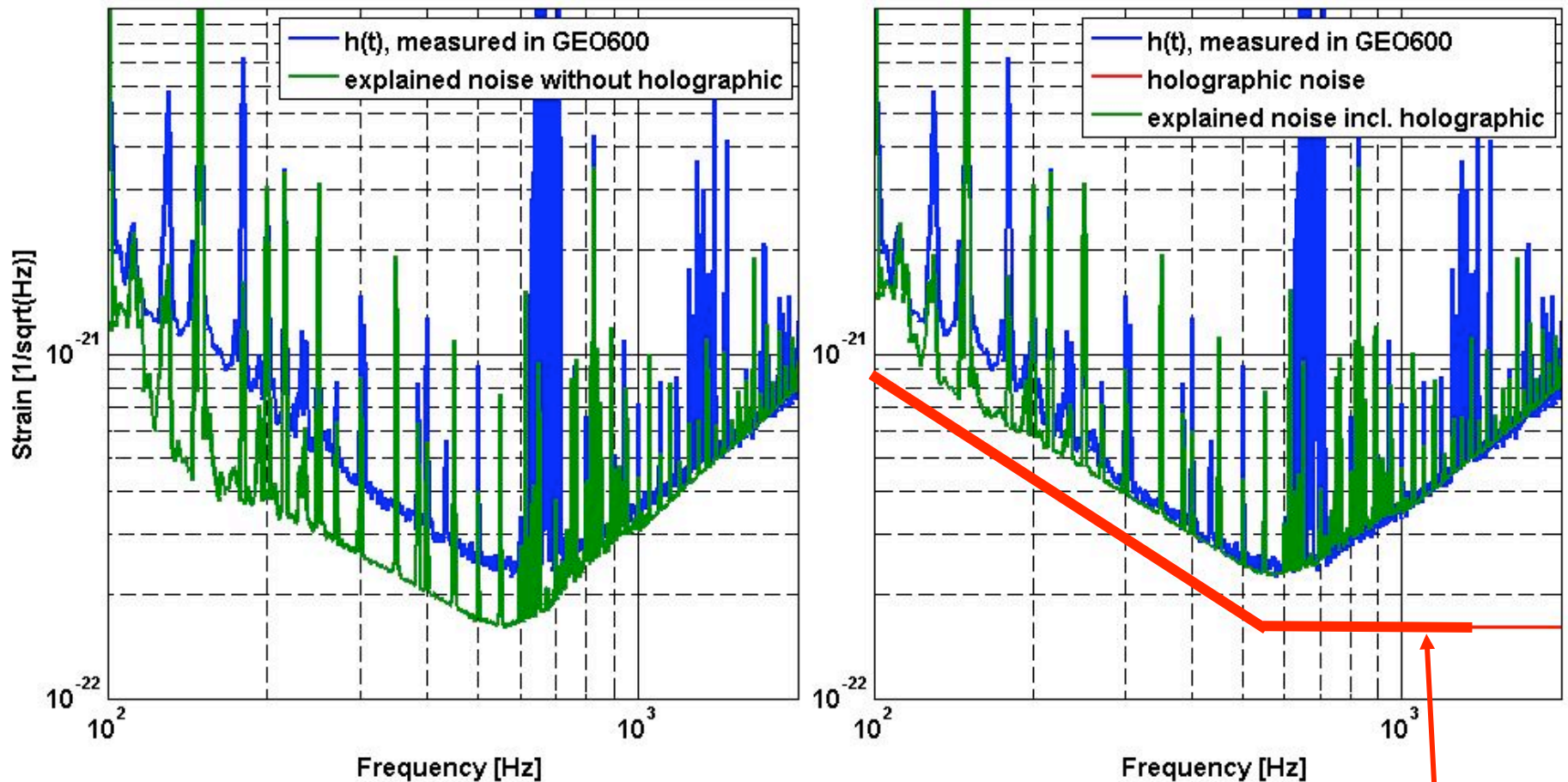


$$h_H(f_0) \simeq \sqrt{t_P/2} = 1.6 \times 10^{-22} \text{ Hz}^{-1/2}$$

S. Hild, GEO600

Frequency (Hz)

# “Mystery Noise” in GEO600



**Data: S. Hild (GEO600)**

**Prediction: [arXiv:0806.0665](https://arxiv.org/abs/0806.0665)**

**Total noise: not fitted**

zero-parameter prediction  
for holographic noise in  
GEO600 ( $\sim$  equivalent  
strain)

## Current experiments

- Most sensitive device, GEO600, sees noise compatible with holographic spacetime indeterminacy
- other explanations, e.g. thermorefractive noise 3.6 times larger than expected in the model, need to be checked
- requires testing and confirmation!
- H. Lück: "...it is way too early to claim we might have seen something."
- LIGO: current data close but may not be sensitive enough, awaits next upgrade
- Models of holographic noise in signals of both systems can be improved

*Interferometers can detect quantum indeterminacy of geometry*

- Beamsplitter inserts holographic uncertainty into signal
- **system with LIGO, GEO600 technology can detect holographic noise if it exists**
- Signatures: spectrum, spatial shear

## New interferometers: beyond GW detectors

- Spectrum:  $\sim 100$  to  $1000$  Hz with existing apparatus
- Higher  $f$  with larger laser power (above GW sources); resonant cavity limit possible
- Test specific geometry dependence (shear character, variation with angle) with different configurations
- Optimal designs different from GW studies
- Different scales: atom interferometers

## Program for holographic noise experiments

- Measure quantum indeterminacy of spacetime
- Measure non-pointlike, wave character of events
- Measure nonlocal quantum weirdness of spacetime metric
- Connect quantum geometry in the lab with information limits from complete fundamental theory (2+1D null projection, black hole entropy, string theory etc.)
- Clues to nature of vacuum-fluctuation gravity, quantum physics of Dark Energy

