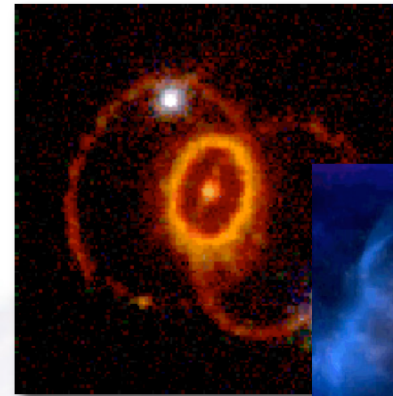


LISA: From the Quantum to the Cosmos

Lee Samuel Finn
Penn State University

Why Gravitational Wave Probes?

- Strong gravitational wave sources are *compact* with internal bulk motion $v \sim c$
- Energetic astronomical phenomena involve strong gravitational potentials
 - $M/r \sim (v/c)^2 \sim 1$
- *Gas, dust accumulate in strong potentials, obscure central engine's EM emission*
- *The most energetic phenomena are EM invisible*



Super-, hyper-novae

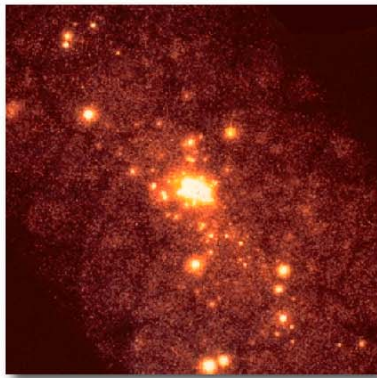
Neutron star coalescence



Black hole binary
coalescence

The universe is transparent to gravitational waves

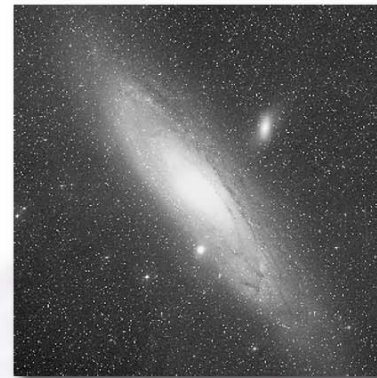
The Universe Today



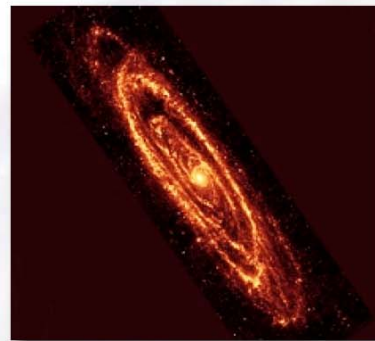
X-ray



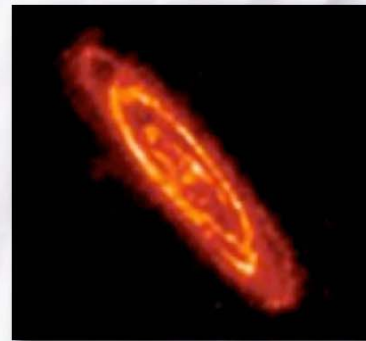
UV



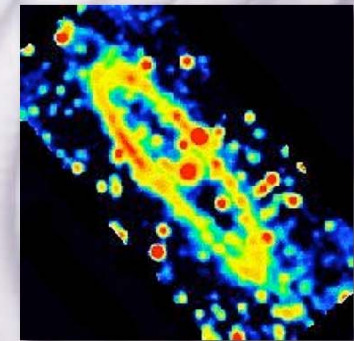
Optical



mid-IR

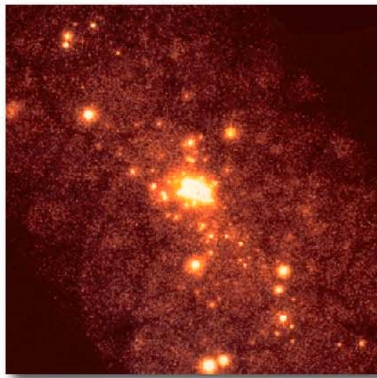


far-IR



Radio

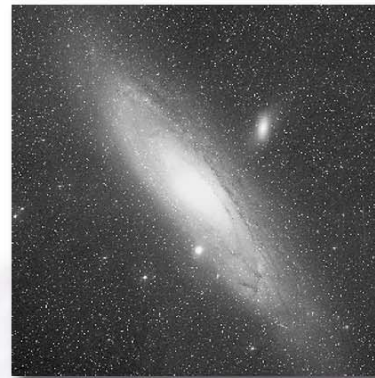
The Universe Tomorrow



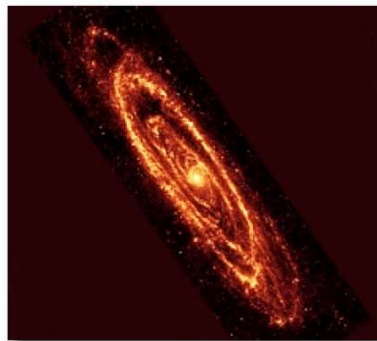
X-ray



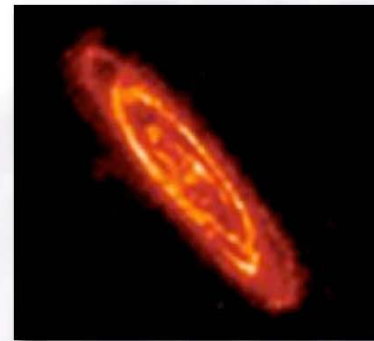
UV



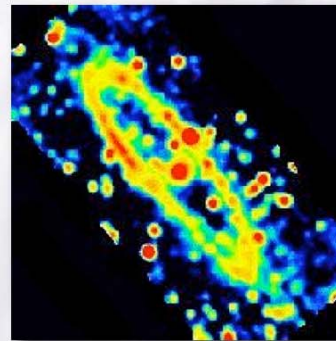
Optical



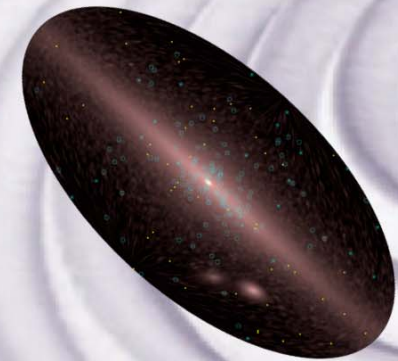
mid-IR



far-IR



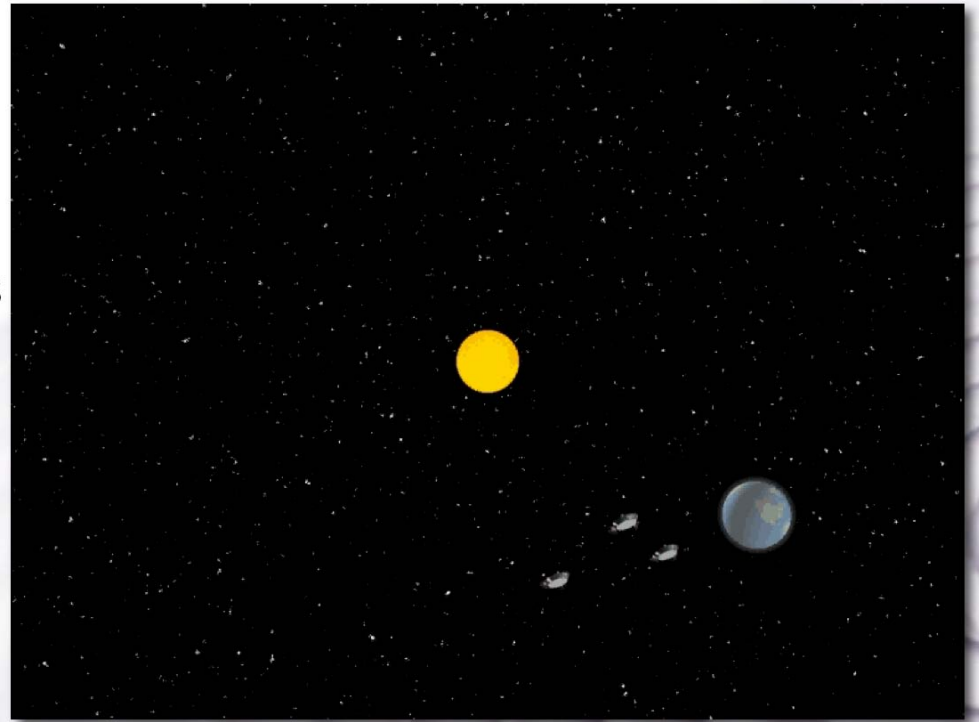
Radio



Grav-Wave

Laser Interferometer Space Antenna

- Three sci-craft constellation in 1 AU circumsolar orbit
 - 5×10^6 km separations
- Measurement: relative velocities to fm/sec precision
 - Grav wave signature: correlated disturbance in relative velocities
- Joint NASA, ESA project
- LISA Pathfinder
 - Technology demonstrator for drag free flight, laser metrology, micropropulsion



Courtesy Rutherford
Appleton Laboratory

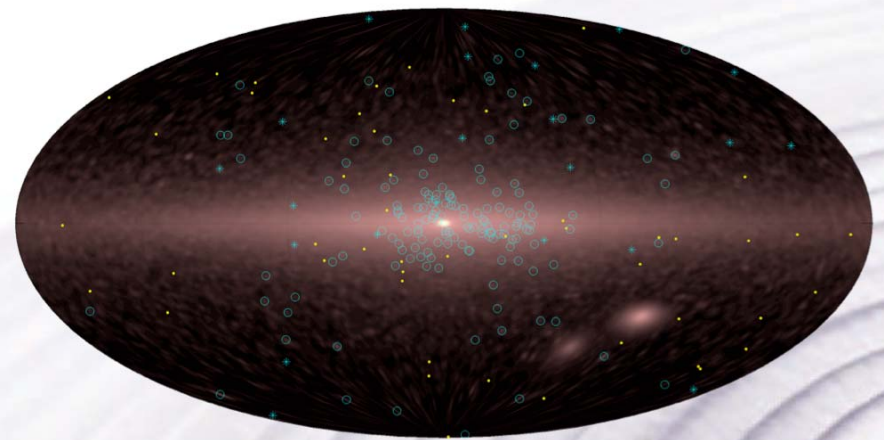
Science Enabled by LISA

LISA's Science Impact is Broad!

- Fundamental Physics of Gravity
- Stellar structure & compact binary evolution
- Galactic structure, observation of “hidden” galactic neighborhood
- Dynamics & dissipative processes in galactic nuclei
- Cosmology & Structure Formation
 - *Absolute calibration of distance ladder lower & upper rungs (galactic & galaxy cluster distances), Galactic nucleus black hole (M , J), merger rates as function of luminosity distance d_L*

Stellar Populations

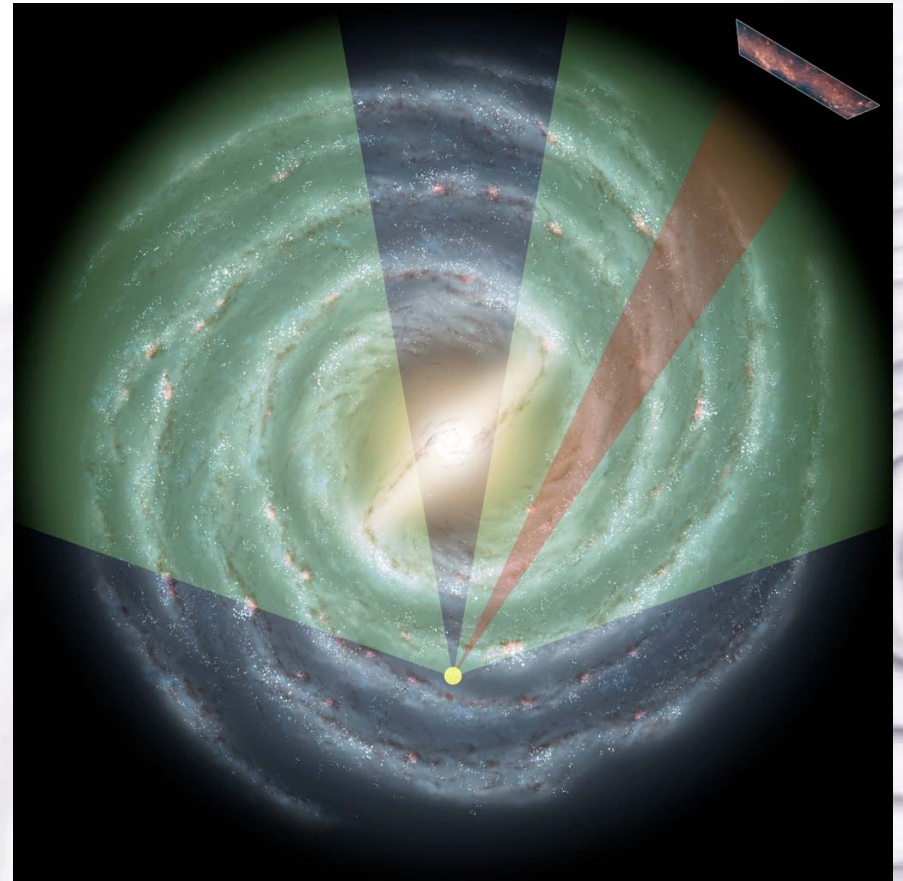
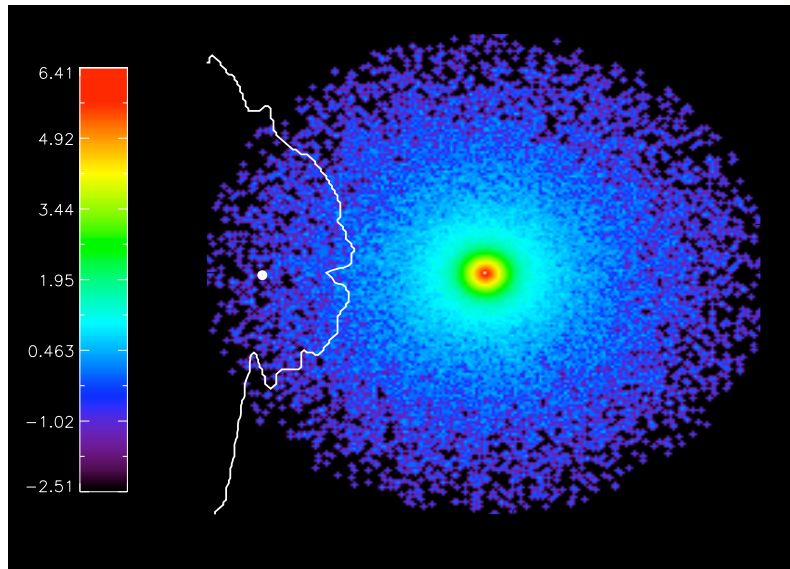
- LISA will identify $\sim 30,000$ galactic WD binaries with $P_{\text{orb}} < 40$ min
 - Determine for each $l, b, P_{\text{orb}}, \sin i, M/d$
- $P_{\text{orb}}\text{-dot}$ measures mass transfer
 - Probe common envelope evolution
- Some fraction identifiable as spectroscopic binaries
 - Determine *full orbital solution, component masses, absolute distance*
- *Calibrate binary stellar synthesis models*



LISA view of the gravitational wave sky
Open circles are globs; yellow dots are (simulated) isotropic sMBH mergers.
Note LMC, SMC in lower-right quadrant.
(Finn, Holley-Bockelmann, Rubbo)

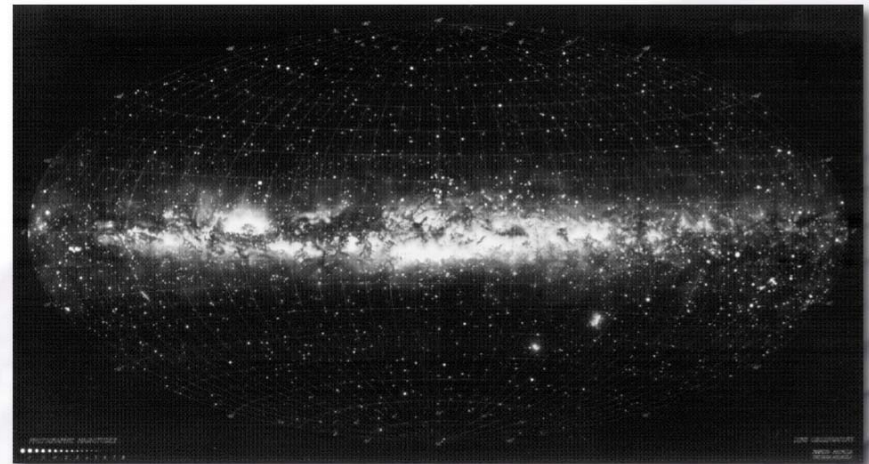
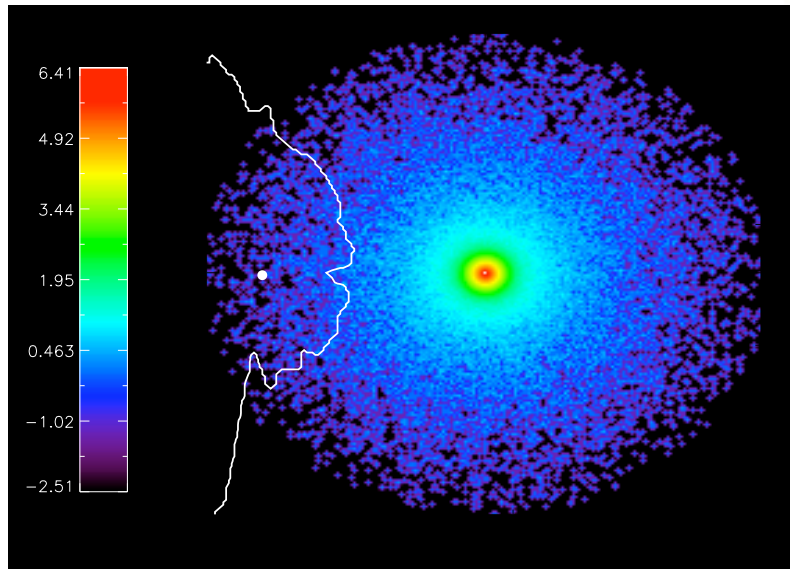
Gravitational waves and galactic structure

Dust, gas, faintness limit EM
observations of galactic
stellar density distribution



Gravitational waves and galactic structure

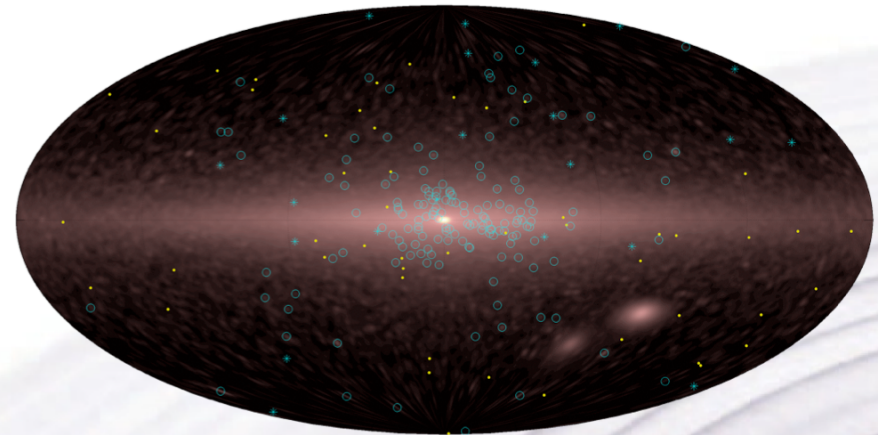
Dust, gas, faintness limit EM
observations of galactic
stellar density distribution



- Barred or unbarred?
- Spiral arms?
- Disk scale height(s)?
- Halo dimensions?

Galactic structure

Galaxy is transparent in gravitational waves, allowing detailed projected imagery of bulge, disk, halo...



LISA resolvable compact binaries

Type	Resolved	With df/dt
(wd, wd)	$>10^4$	~ 600
AM CVn	$>10^4$	~ 50
(ns, wd)	21	3
Other	2	0

Nelemans 2003

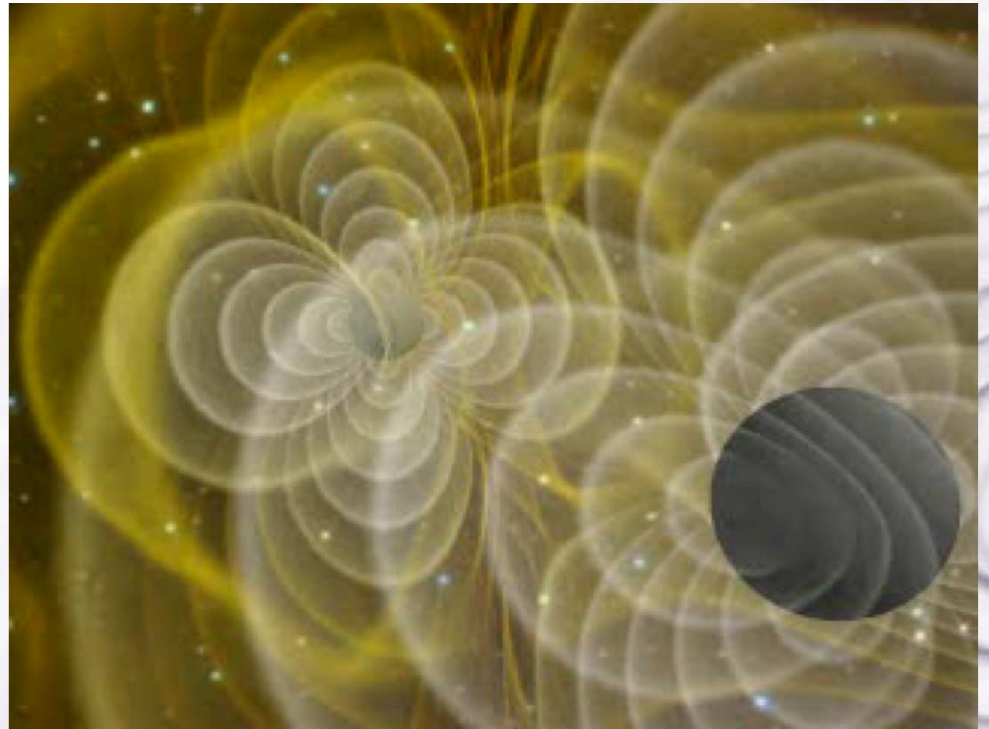
- Bar dimensions, orientation, disk & halo scales, spiral arms, etc.
- “Zone of avoidance” dwarf/satellite galaxies, globular clusters...
- Binary mass function from consistency with galactic model...

Extra-galactic astronomy

- Massive black holes are ubiquitous in galactic centers
 - At least, we see evidence for them wherever we are able to ...
- Every galaxy has undergone a major merger in its past
 - Black holes sink to potential center and coalesce
 - *Dissipative processes by which black holes traverse final parsec are not understood*
- LISA can localize BH mergers to cluster (& possibly galaxy) scale
 - *Morphology at merger measures dissipative process timescales*



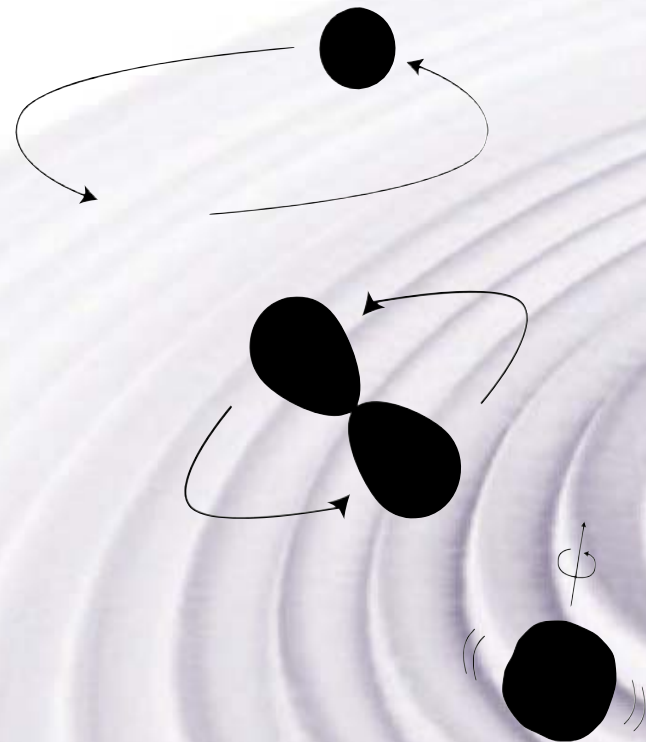
Measuring Black Hole Hair



Credit: GSFC Gravitational Physics Lab

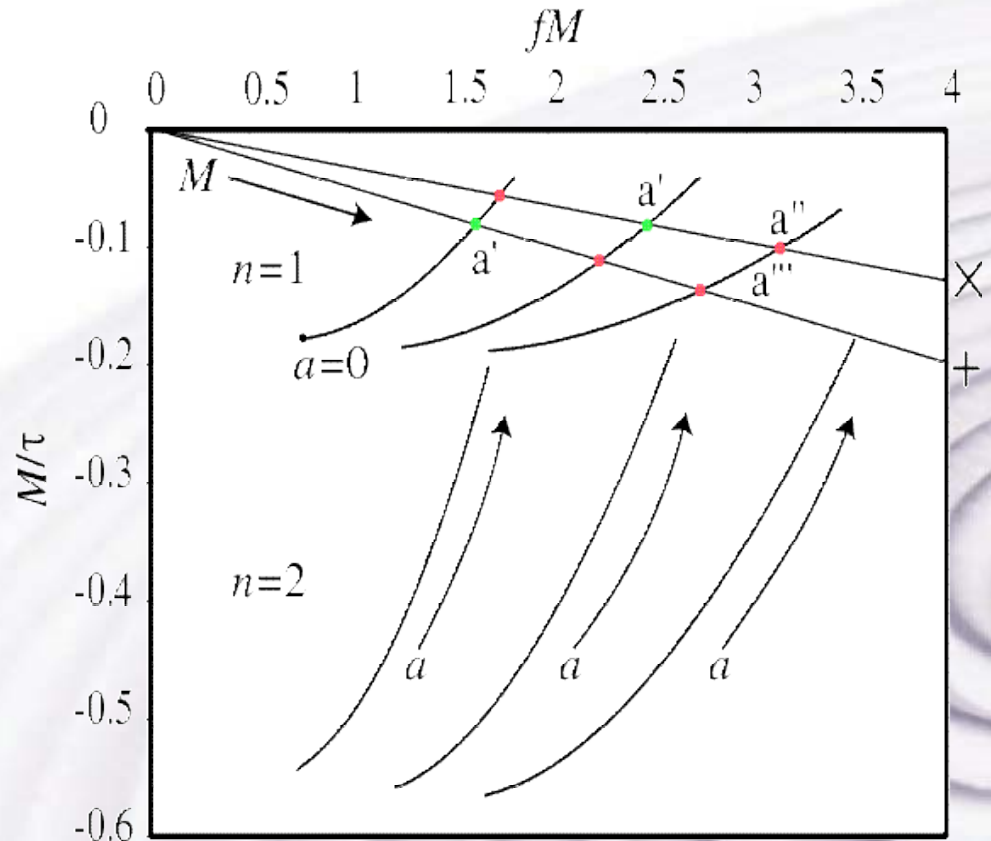
Measuring Black Hole Hair

- What distinguishes between black holes? Hair!
 - “Ringdown”: damped sinusoidal signal
 - $s(t) \sim \sum \exp(-\pi f_k t / Q_k) \sin 2\pi f_k t$
 - Einstein: Ringdown f_k, Q_k depend uniquely on $M, J/M^2$
 - *Other length scales will disturb relationship*



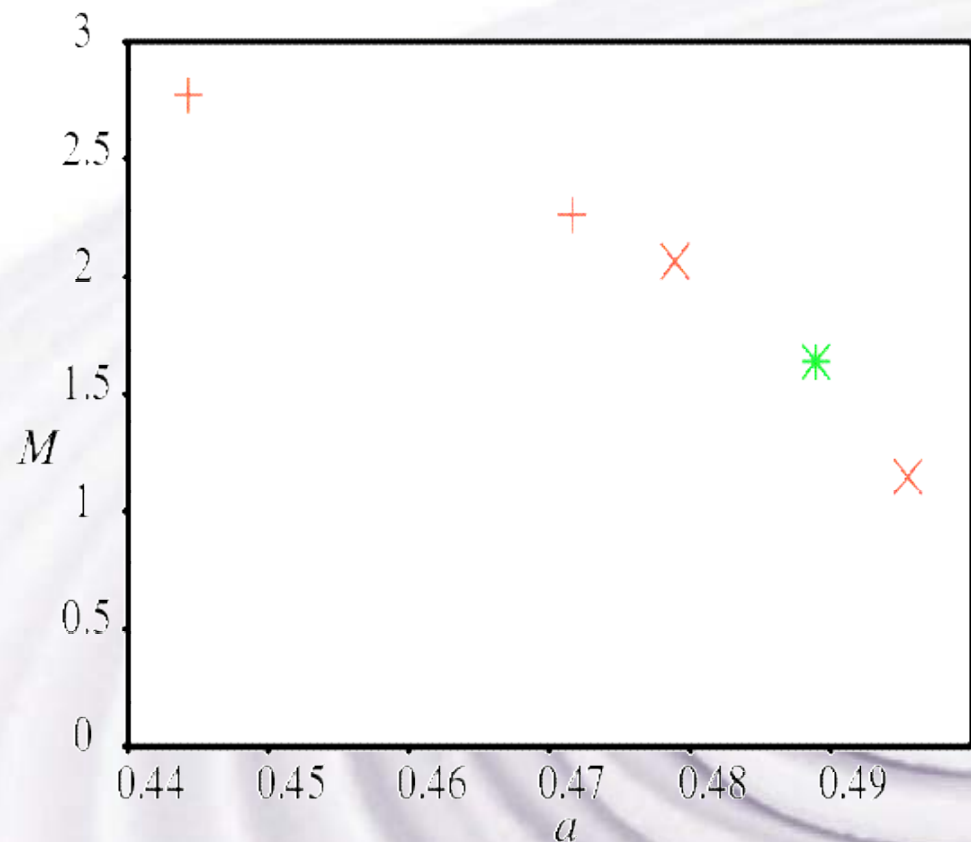
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- Observe (f, Q) pairs
 - Each pair suggests set of (M, a, n, l, m) n -tuples

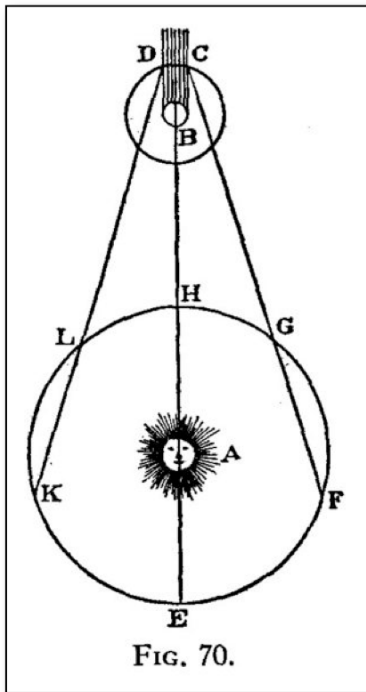


Measuring Black Hole Hair

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- *Other length scales will disturb relationship*
- Observe (f, Q) pairs
 - Each pair suggests set of (M, a, n, l, m) n -tuples
- “No-hair” theorem: “Einstein consistent” (in $M, J/M^2$) f, Q pairs
 - Deviations are evidence for other length scales



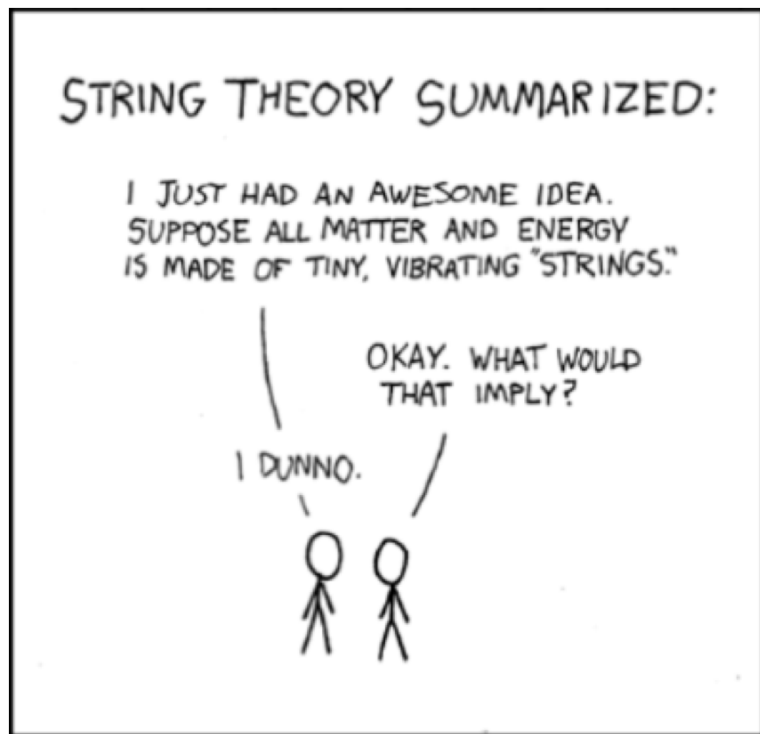
“Demonstration touchant le mouvement de *la attraction universelle*”



- 1676: Roemer measures c from “Galilean Clock” arrival time variations over Jovian synoptic year
- GW observation of a pulsar, or binary measures c_g :
 - $c_g = \frac{8\pi}{\Delta\Phi_{PP}} \frac{R \sin \theta}{P_{NS}}$
 - LISA + WD Binary: $\frac{\Delta c_g}{c} < 10^{-3}$
 - LIGO + Pulsar: $\frac{\Delta c_g}{c} < 10^{-6}$

cf. Roemer *Journal des Sçavans* 7 Dec 1676, pg 233-6;
Finn & Romano, Woan & Hendry, in preparation

String theory and quantum gravity



- Consistency of string theory, standard model extension gravity require GR quantum correction
 - Chern-Simons term *affects only gravitation sector*
 - Affects propagation: a wavenumber dependent *birefringence* amplifies one (circular) polarization, suppresses other
- Wavenumber of $\sim 10^6 M_{\odot}$ inspiraling binary signal changes adiabatically by $\times 100$ over observation
 - "Apparent" anomalous binary precession
- LISA can observe cosmological binary coalescences (z up to 30!)
 - Bound/measure *form-factors* associated with cosmologically averaged CS correction!

LISA

- Exceptionally broad Science Impact & Discovery Potential
- Truly from the Quantum to the Cosmos!
- Observations enable scientific investigations not otherwise possible
- Pathfinder verifies critical technologies
- Superior discovery potential in frontier science for the next decade!

