

PMT Calibration School

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PMT Calibration School



Please take a lecture note!

PMT Calibration School

Saturday, April 26

10:00	12:30	Basics of PMT Calibration (K. Arisaka, Instructor) <i>ROOM MAITEN (BLUE)</i> <u>Principle of PMT and its calibration (1hour)</u> <u>Absolute Calibration by NIST Standards (30min)</u> <u>Imperfect behavior of PMT (30min)</u>
12:30	13:30	Visit to PMT test/potting facility at SDE-Fabrica (Shuttle bus leaves the Conference center at 12:30)
15:00	18:00	SD and FD PMT Calibration (K. Arisaka, Instructor) <i>ROOM CANELO (RED)</i> <u>SD PMT Specs and Calibration (30min)</u> <u>SD PMT Test Facility and Results (30min, Barnhill)</u> <u>FD PMT Specs and Calibration (20min)</u> <u>FD Photon Yield and Systematic Errors (50min)</u>

SD Calibration Workshop

Sunday, April 27

10:00	13:00	SD Calibration Workshop (A. Tripathi organizer) <i>ROOM MAITEN (BLUE)</i> Local Station Calibration and Monitoring Update: Xavier (5min) Studies of Pedestal Induced Systematics and Electronics Noise: Gonzalo (20min) Report from Puebla: Humberto (20min) LED Flasher Measurements in Orsay Tank, and Result of Biological analysis of EA Tanks: Carla (25min) Water Level Monitoring Studies: Clementina (20min) Results of the muon data Analysis: Arun (30min) The Offline Calibration Database: Tom Paul (15min) Thoughts on SD Calibration in Pre-production: Xavier (20min)
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SD/FD Cross Calibration Workshop

Sunday, April 27

14:30	18:00	FD/SD Energy Cross Calibration (K.Arisaka, organizer) <i>ROOM MAINTEN (BLUE)</i> <u>Absolute Energy by SD: AGASA/Auger-SD</u> (1hour) Auger FD Dome Calibration (20min, Jeff Brack) Auger FD Laser Calibration (20min, Mike Roberts) <u>Absolute Energy by FD: HiRes/Auger-FD</u> (50min) <u>How to Minimize Systematic Error in FD</u> (30min) <u>Strategy for SD/FD Energy Cross Calibration</u> (30min) Discussion (All, 30min)
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Homepage of PMT School

➤ **Homepage is**

http://www.physics.ucla.edu/~arisaka/auget/pmt_school/

➤ **All the presentations are available.**

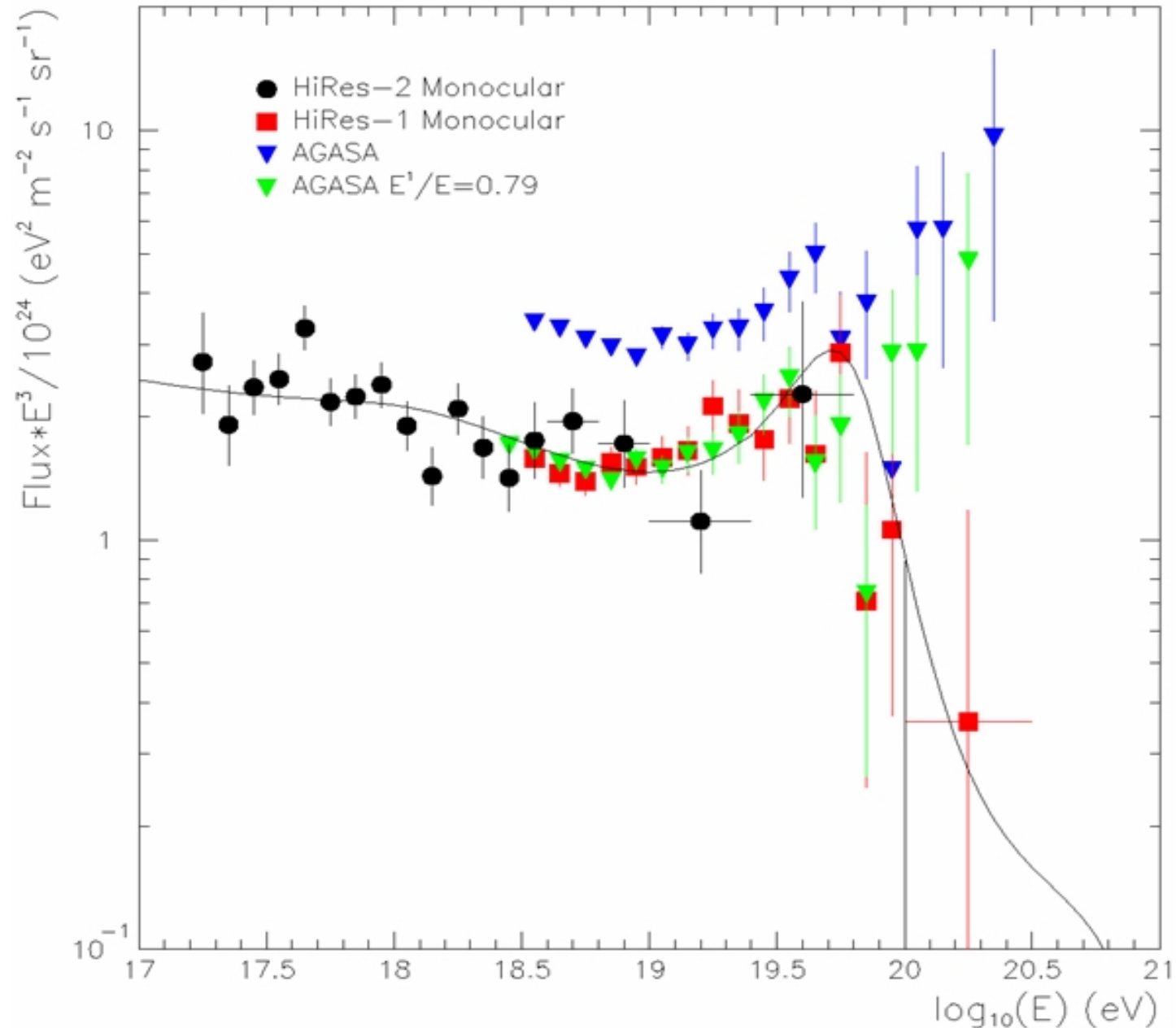
- **PPT files**
- **PDF files**
- **Combined PDF files (6 pages in 1 page)**

Expected Attendees

- **Active workers at UCLA/UTN PMT test facility at Malargue.**
 - **Local group.**
 - **Volunteers from outside.**
- **Whoever interested in “absolute energy” of UHECR.**
 - **Experts in SD and FD detector calibration.**
 - **Experts in SD and FD data analysis.**

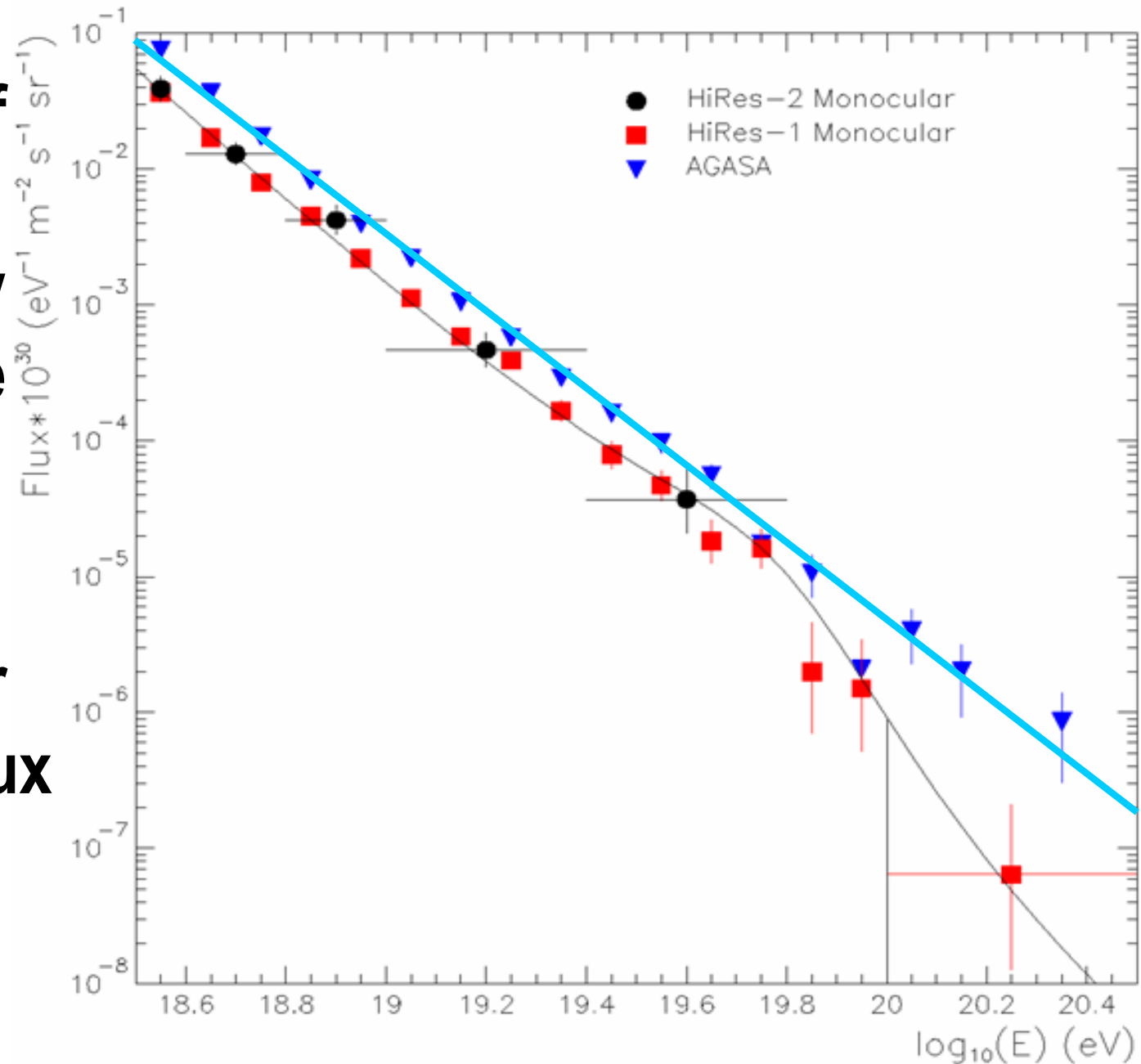
What is the Problem?

- AGASA sees a dozen of super GZK Events.
- But HiRes sees GZK cut-off?
- A factor of two difference in the Flux below GZK cutoff.



HiRes vs. AGASA

- Flux difference of the factor of two can be explained by energy difference of 20-30%.
- Is there any systematic error contributing to flux distortion above GZK cutoff?



What are we facing?

➤ Is the Pierre-Auger going to uncover their mistakes and fix them?

or

➤ Are we going to repeat the same mistakes?

➤ That is the question of systematics, not statistics.

Striking Fact

- If Hamamatsu miss-measured the Quantum Efficiency and/or Collection Efficiency of “one particular PMT” used by Kakimoto (say, it was 20% too low), it can explain the inconsistency between AGASA and HiRes.
- Moreover, Auger-FD would repeat the same, wrong spectrum of HiRes.

Why do we have to care PMTs?

- **In FD, the absolute energy scale comes from**
 - **PMT calibration in photon yield study**
 - **PMT calibration in telescopes**
- **In both SD and FD, imperfect behavior of PMTs is the source of systematic errors**
 - **Stability**
 - **Non linearity**
 - **After pulse**
 - **...**

Goal of PMT School

- **Understand Basic Principle of PMT.**
- **Become familiar with measuring various PMT parameters.**
- **Find out possible systematic errors in previous experiments.**
- **Try to minimize them in Pierre-Auger experiment.**

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